



# Spatial photon correlations using nearly dead time free ultra-high throughput single photon detection

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Quantum Optics and Quantum Information

Friedrich-Alexander Universität Erlangen-Nürnberg

Stellar Intensity Interferometry Workshop 2024



# Agenda

1. Introduction
2. Measurement Setup
3. Temporal Correlations at C2PU
4. The Problem of moving Stars
5. Spatial Correlations using HPDs
6. High Throughput System (HTS)
7. Spatial Correlations using HTS
8. Summary and Outlook

2023



January						
S	M	T	W	T	F	S
1	2	3	4	5	6	7
8	9	10	11	12	13	14
15	16	17	18	19	20	21
22	23	24	25	26	27	28
29	30	31				

February						
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	1	2	3	4		
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12	13	14	15	16	17	18
19	20	21	22	23	24	25
26	27	28				

March						
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26	27	28	29	30	31	

April						
S	M	T	W	T	F	S
						1
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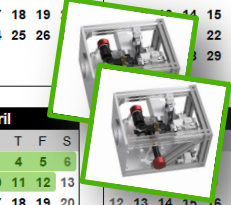
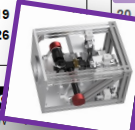
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21	22	23	24	25	26	27
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June						
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						1
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2024

July						
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August						
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16	17	18	19	20	21	22
23	24	25	26	27	28	29
30	31					



September						
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						1
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October						
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30	31					

November						
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						1
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23	24	25	26	27	28	29
30						

<https://www.vertex42.com/calendars/2023.html>

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January						
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30	31					

February						
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March						
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April						
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23	24	25	26	27	28	29
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May						
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July						
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August						
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16	17	18	19	20	21	22
23	24	25	26	27	28	29
30						

December						
S	M	T	W	T	F	S
						1
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9	10	11	12	13	14	15
16	17	18	19	20	21	22
23	24	25	26	27	28	29
30	31					

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# Introduction

# Motivation

- Larger telescopes increase SNR in HBT measurements
- For bright stars: hybrid single photon detectors saturate for telescopes  $> 1\text{m}$
- For spatial correlations ordered timestream is necessary
- Manageable data rate



## Solutions:

- 1) Use new kind of detector
- 2) Find Synchronization hardware
- 3) Use high throughput TDC



# Telescope Site

The Calern observatory is located in the South of France close to Nice on a plateau with an altitude of 1270m. C2PU offers twin telescopes at a separation of 15m.



- Primary mirror diameter of 1.04 m
- Equatorial yoke mount
- Cassegrain secondary focus
- Focal length of 13m and F/12.5



06°55'23" East  
43°45'13" North



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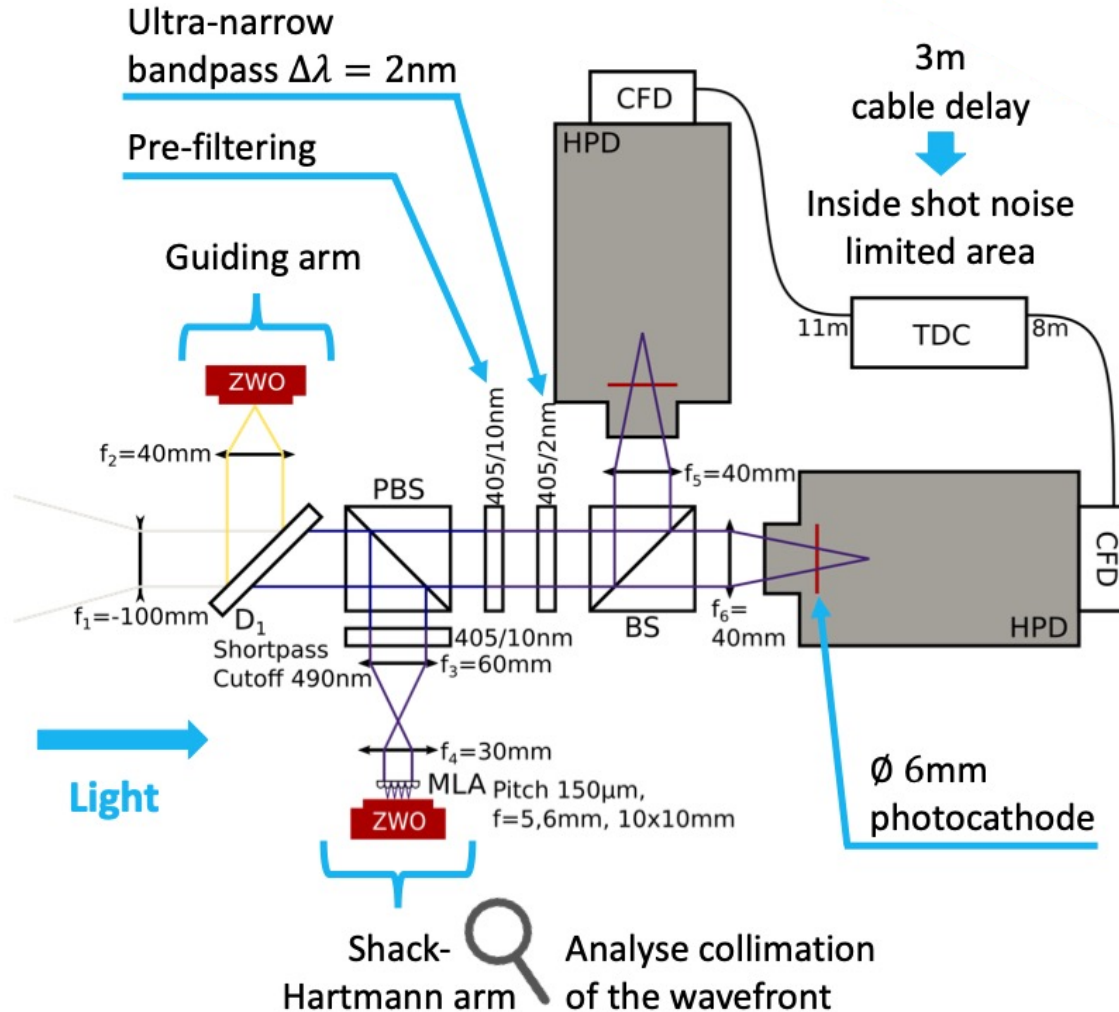
East Dome  
Epsilon  
Omicron  
West Dome





# Measurement Setup

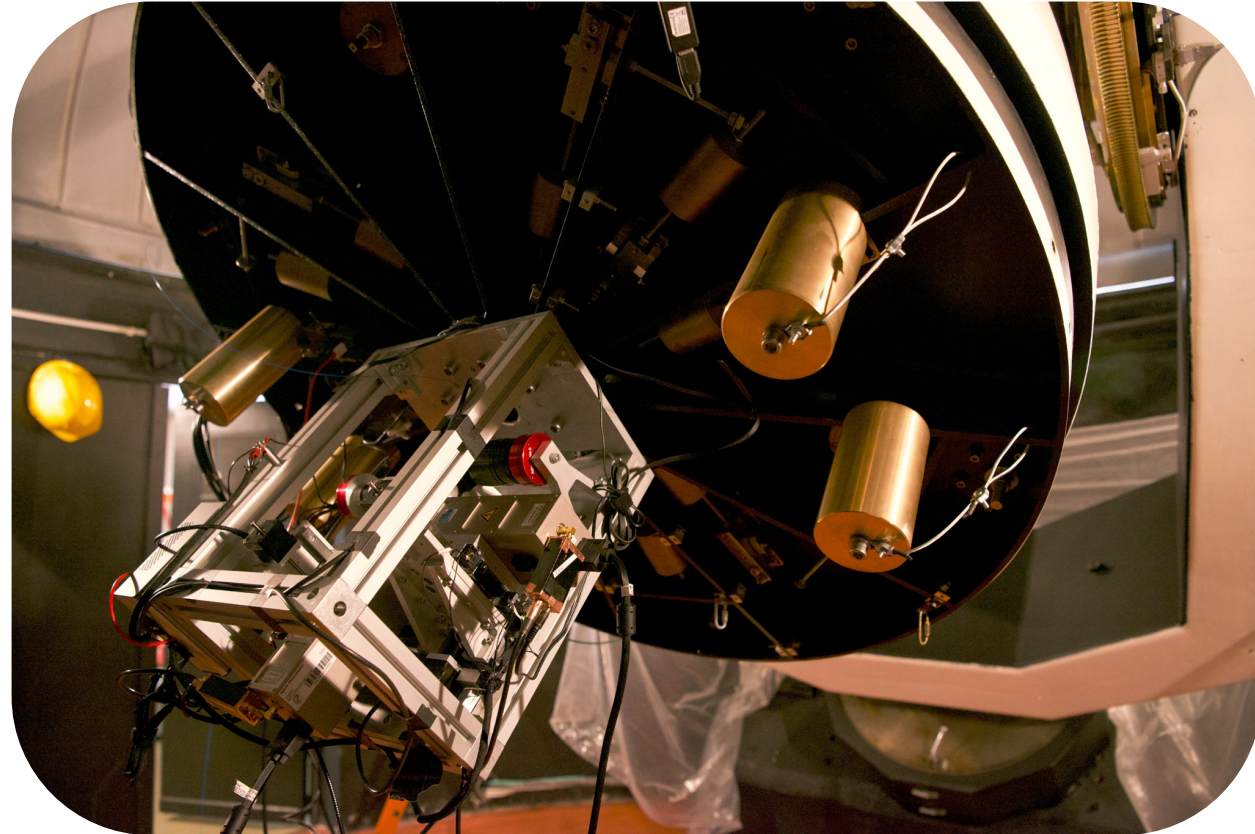
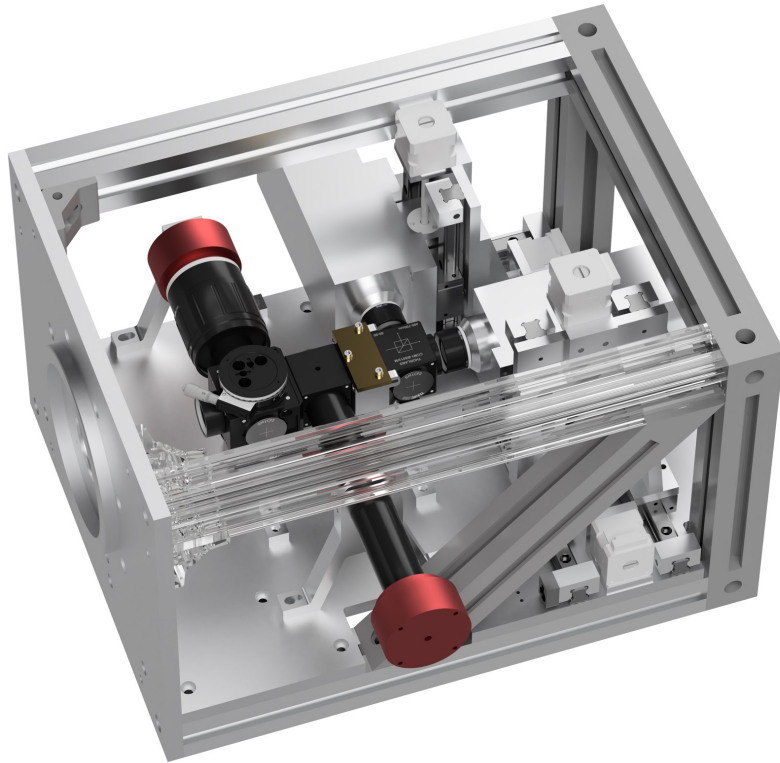
# Optical Setup



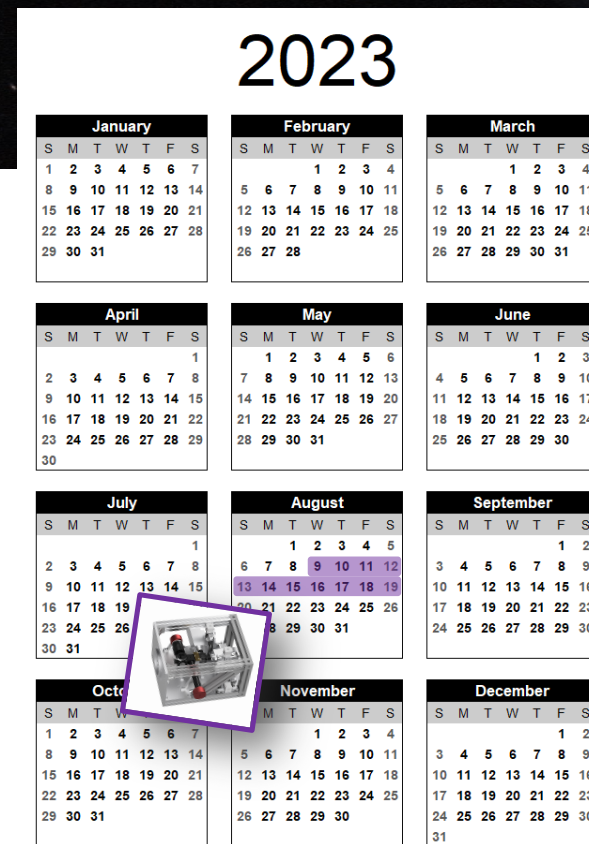
- 40/45mm lenses for beam compression (no focus)
- Used for all three measurements
- TDC and detectors will change
- Cable delays will vary from 3 to 12m



# Measurement Setup

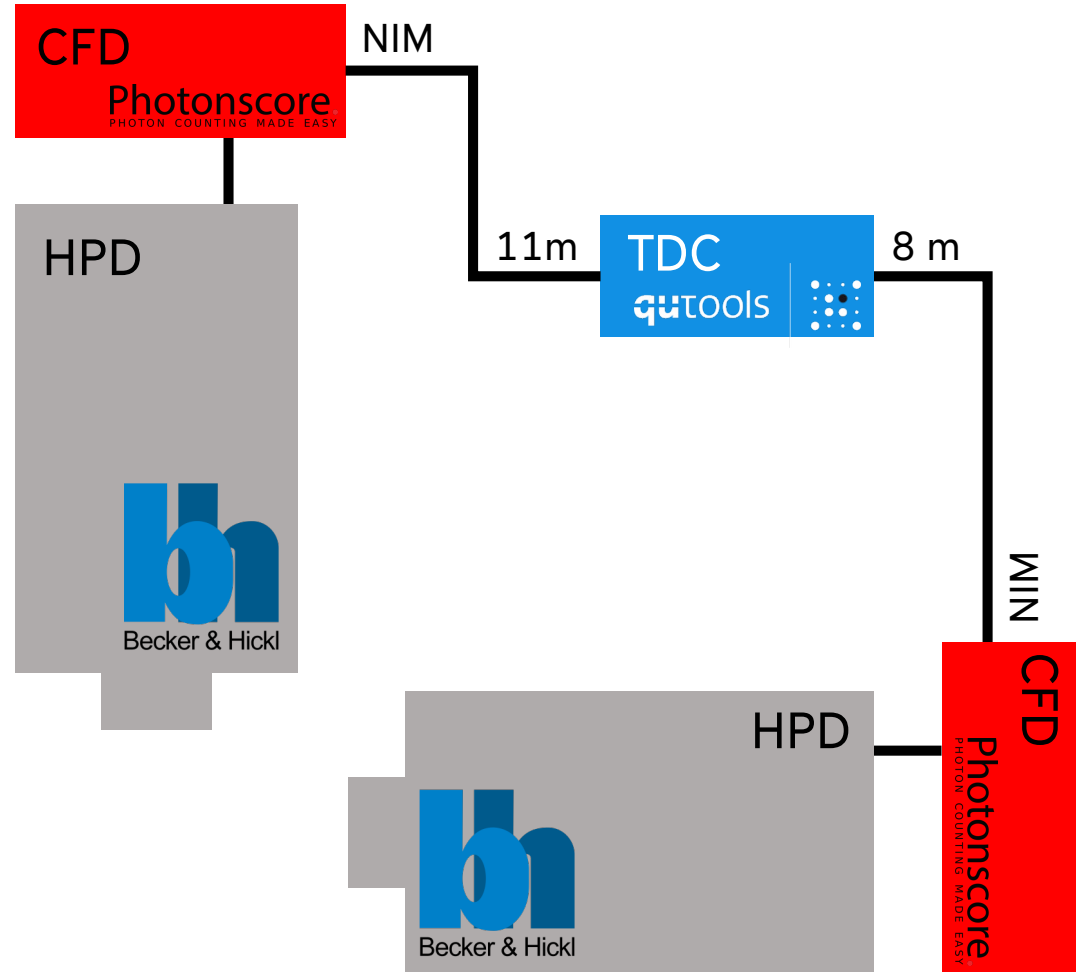
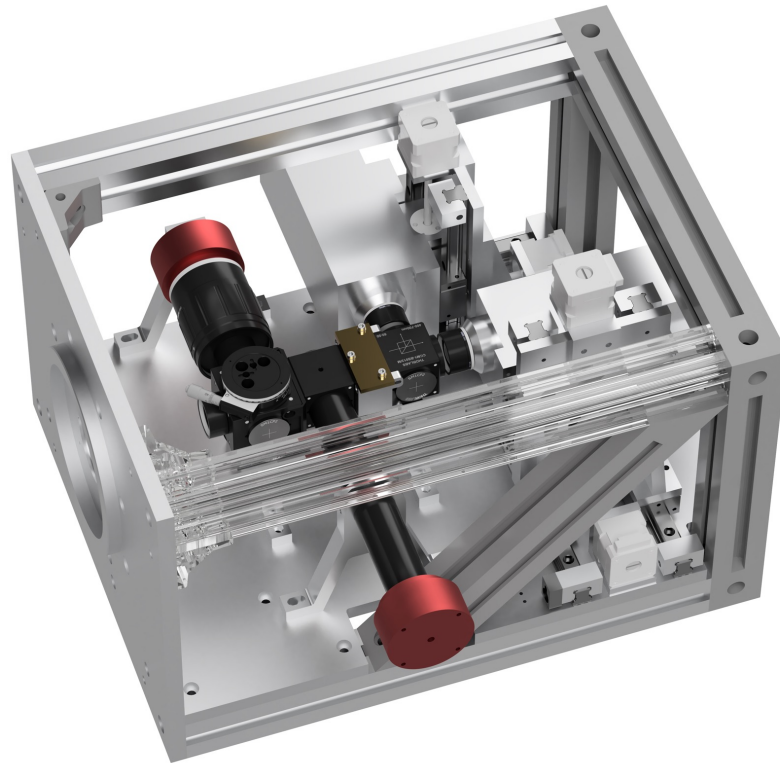


# Temporal Correlations at C2PU





# Measurement Configuration



# Hybrid Photon Detectors (HPDs)

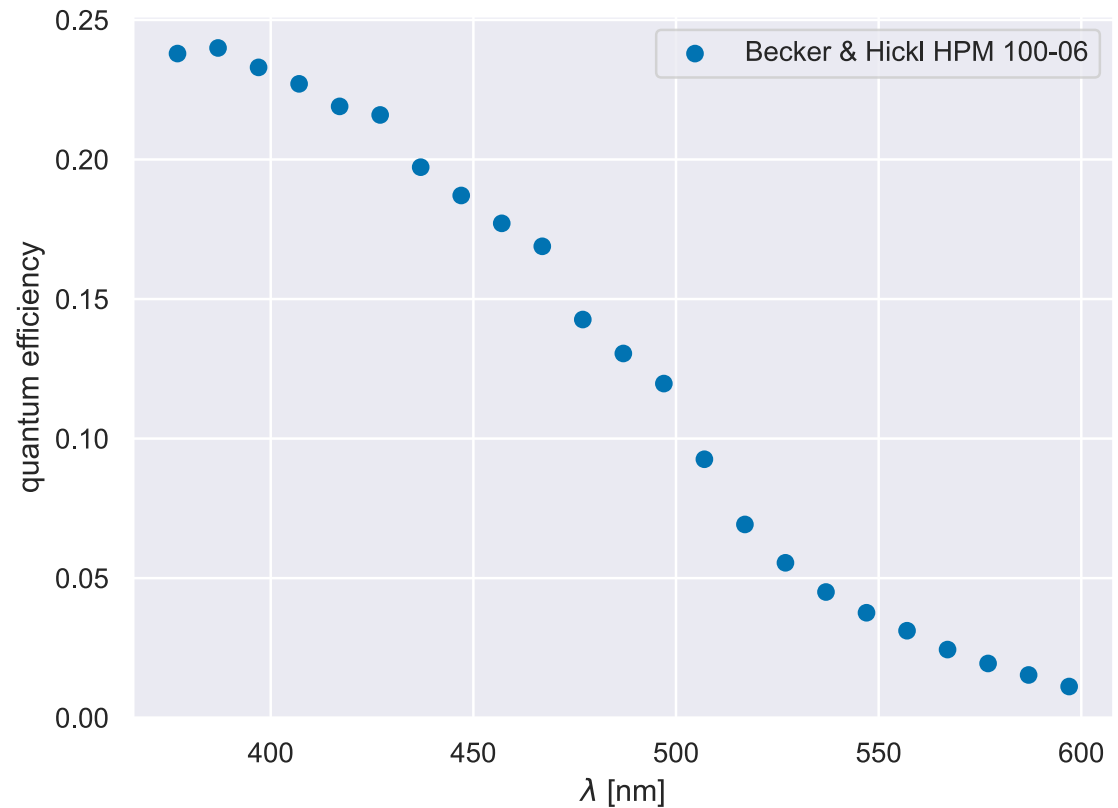
- HPM-100-06 from Becker&Hickl
- Hamamatsu R10467 hybrid detector tube
- $\varnothing 6\text{mm}$  active area
- Max. continuous countrate 10MHz
- 100 - 400Hz dark counts
- Quantum efficiency of 22.7% at 405nm
- Output is not discriminated, puls height varies



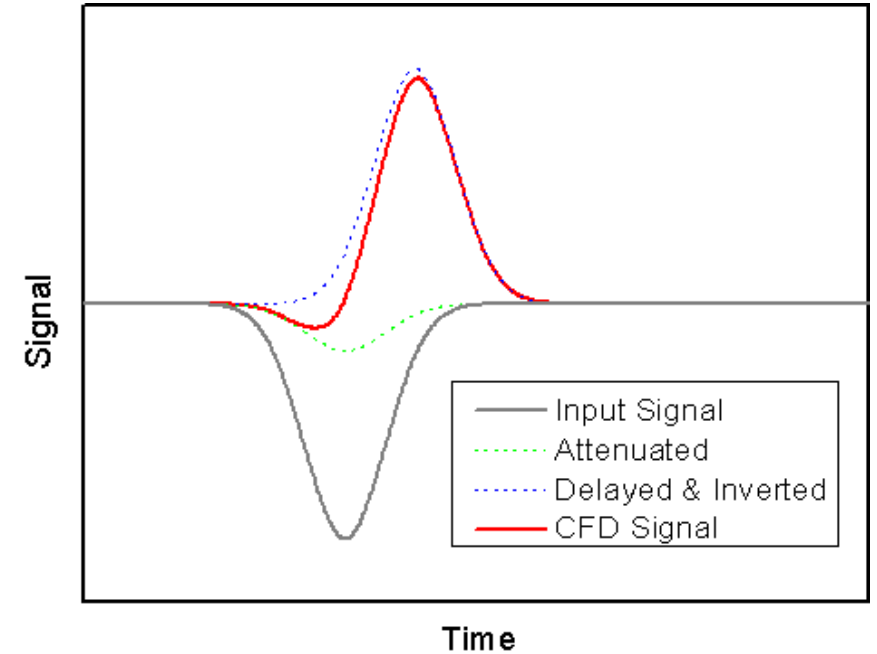
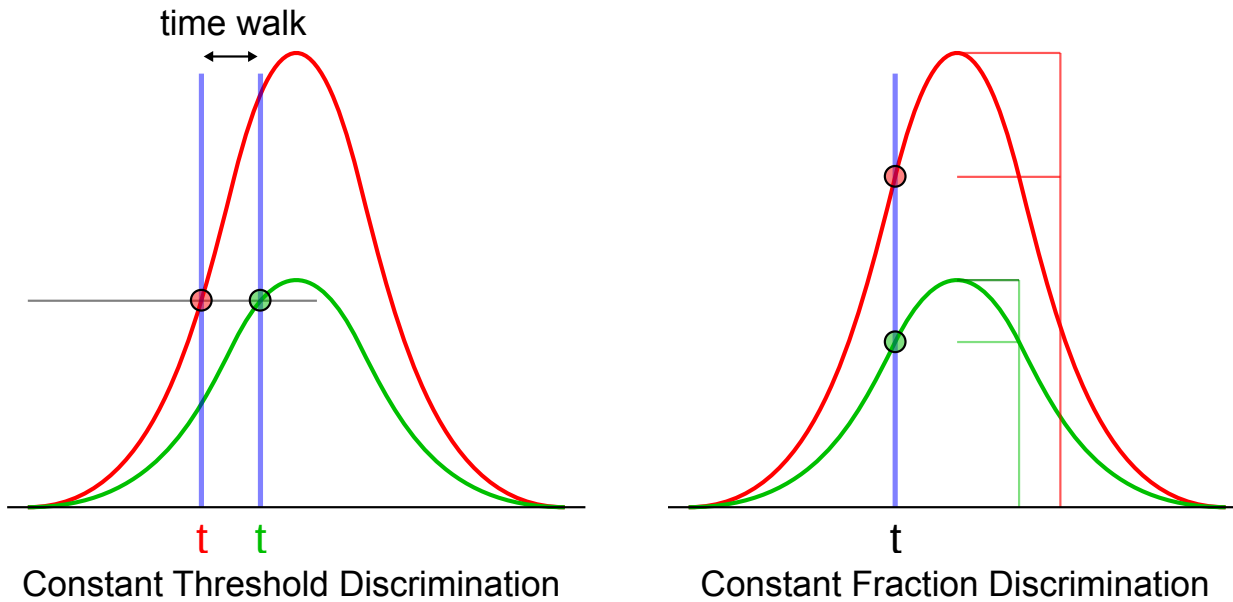


# Hybrid Photon Detectors (HPDs)

- HPM-100-06 from Becker&Hickl
- Hamamatsu R10467 hybrid detector tube
- $\emptyset$ 6mm active area
- Max. continuous countrate 10MHz
- 100 - 400Hz dark counts
- Quantum efficiency of 22.7% at 405nm
- Output is not discriminated, puls height varies



# Constant Fraction Discriminator (CFD) FAU



→ Use Constant Fraction instead of Constant Threshold Discriminator

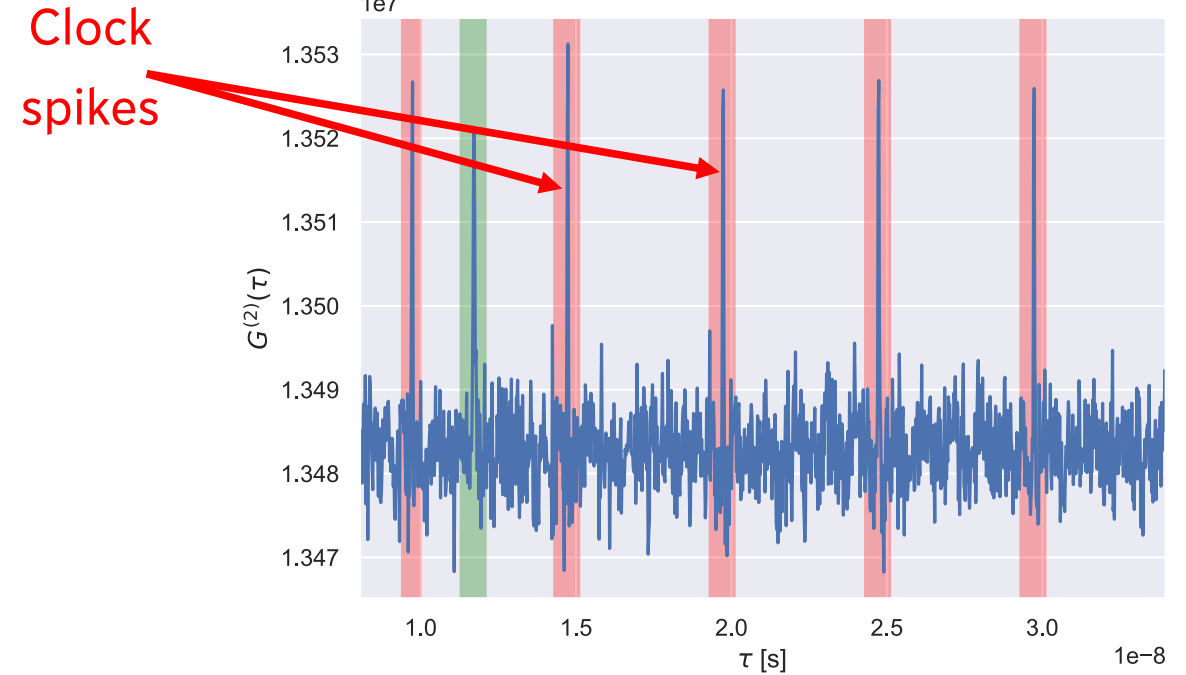
→ Tune parameters *threshold* and *zero crossing*



# quTools TDC



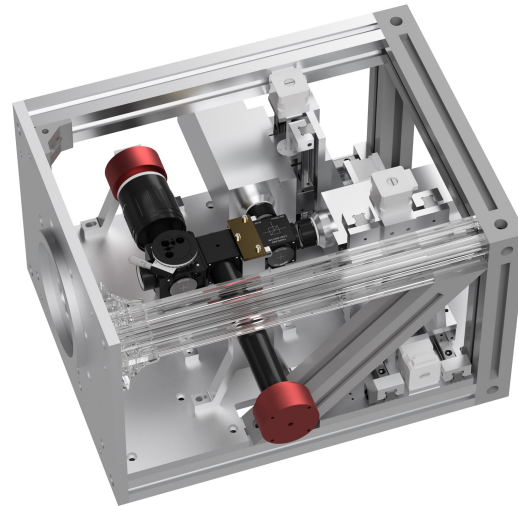
- 40ns deadtime
- maximal event rate per channel is limited to 25 Mcps



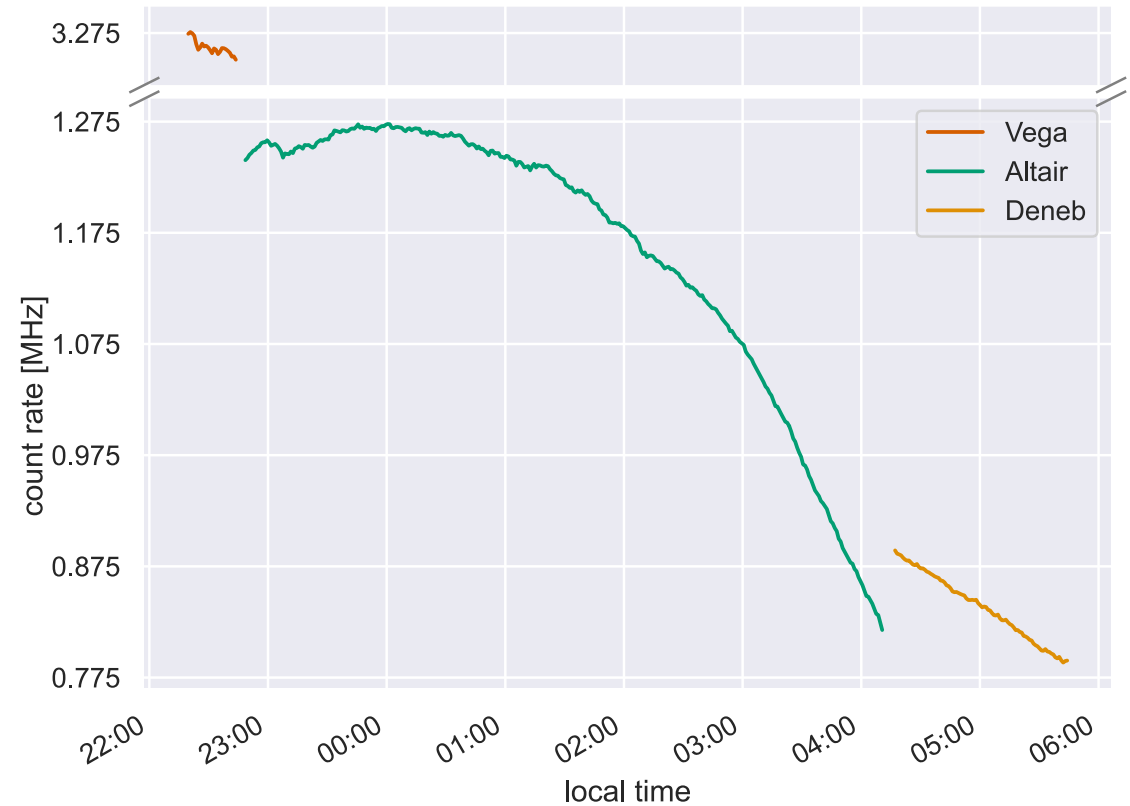
→ 3m cable difference needed to lie inside a shot noise limited area

# Advantages of the Setup

- Tracking of the star never lost
- Short optical path
- Observation in the blue at 405nm
- Easy to align at the telescope
- Movable detectors in x and y-direction



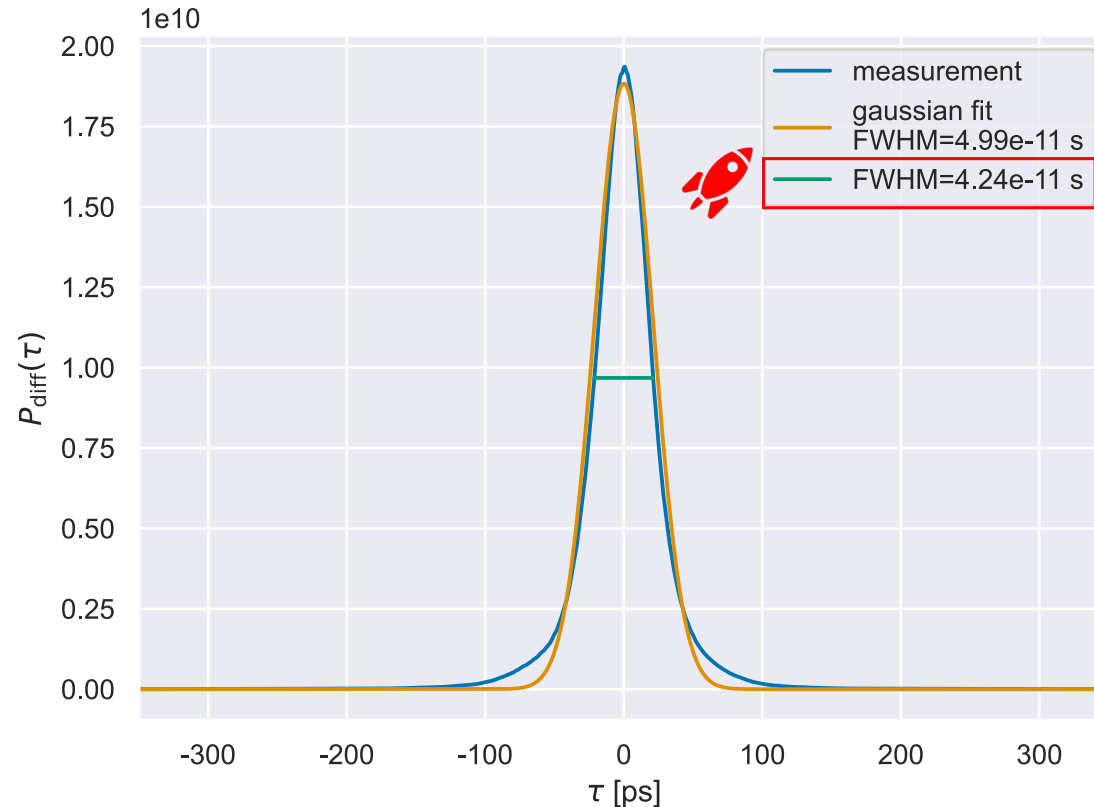
## Typical Measurement Night





# Timing resolution of the setup

Timing jitter @2MHz



- **42.4ps timing jitter** of the system using a fs-pulsed laser
- Full photo detection system of 2 HPDs, 2 CFDs and the quTAG, as well as one 50:50 non-polarizing beam splitter
- Result will be used to calculate the measurement expectations

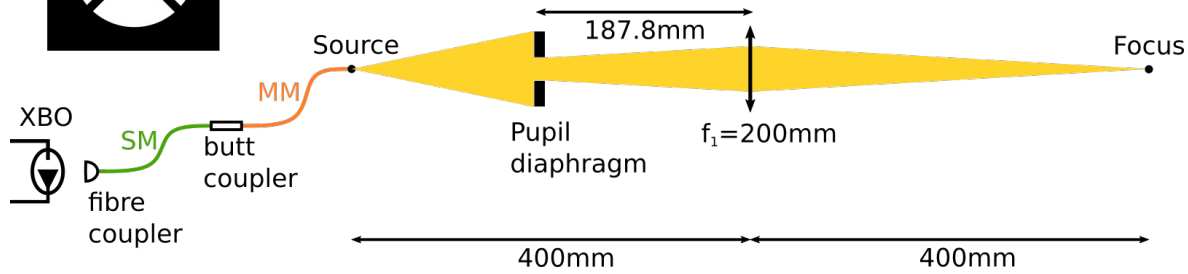
# Observe artificial star

Pupil diaphragm

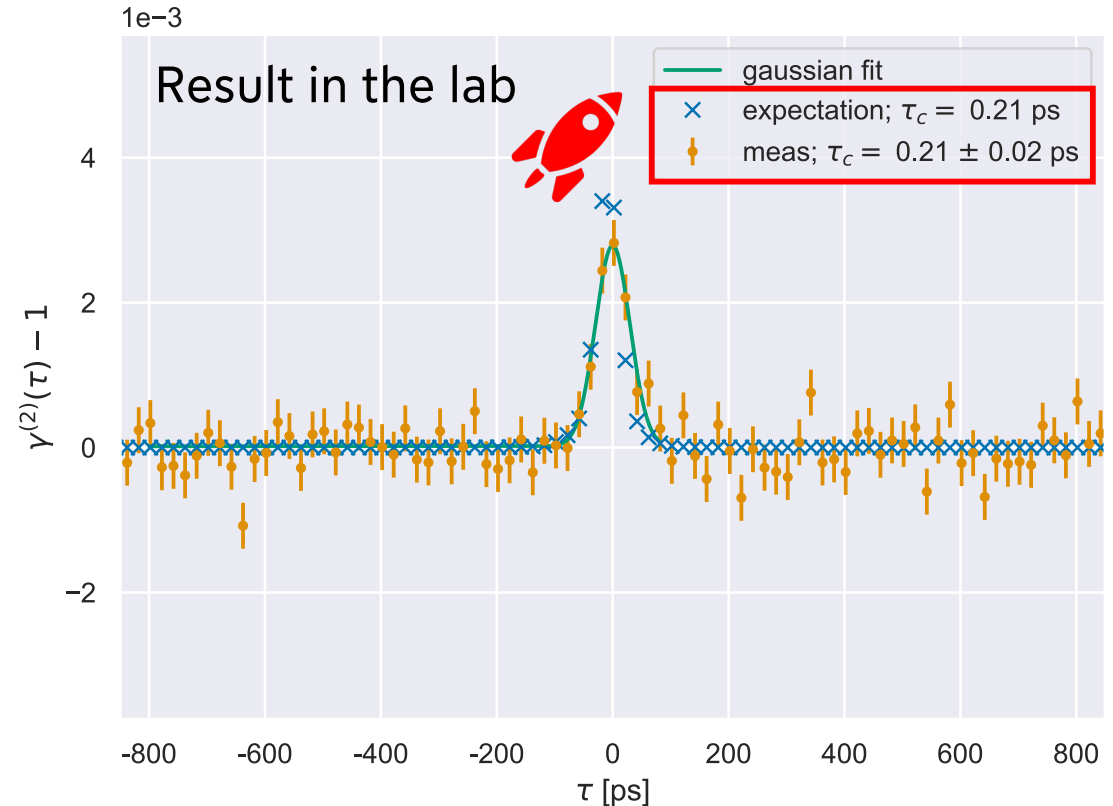
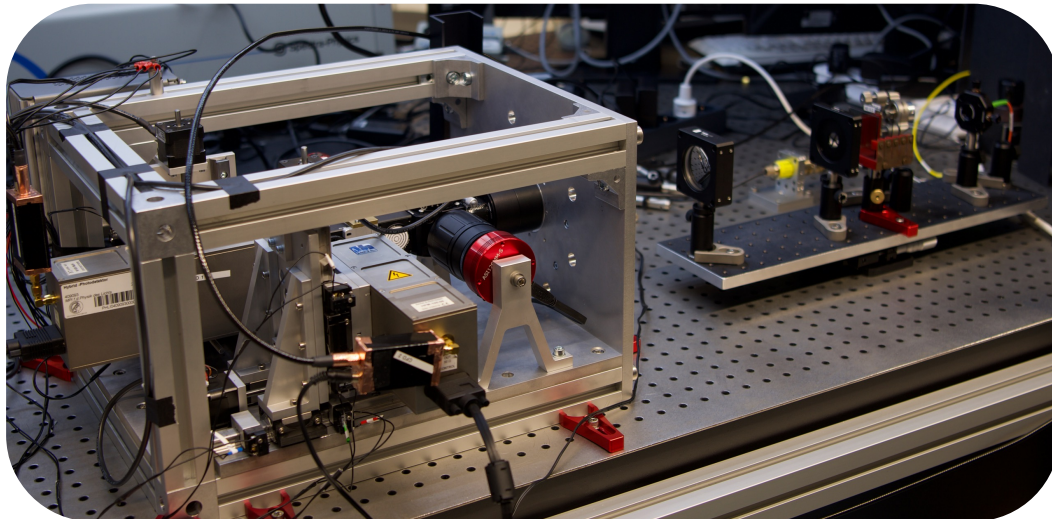


$d=4,9\text{mm}$   
 $D=17\text{mm}$

2f - 2f configuration



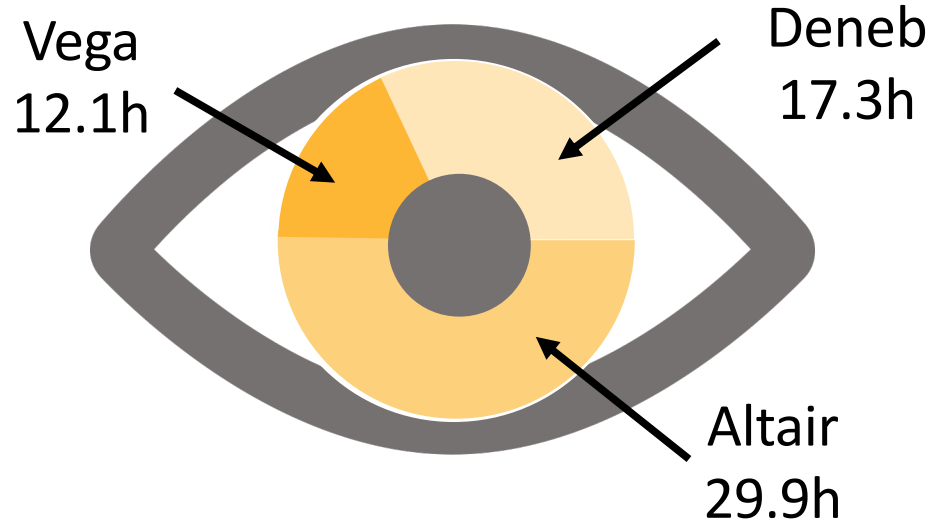
- Expected coherence time **0.21ps**
- Measured coherence time with artificial star in the lab:  **$0.21 \pm 0.02\text{ps}$**



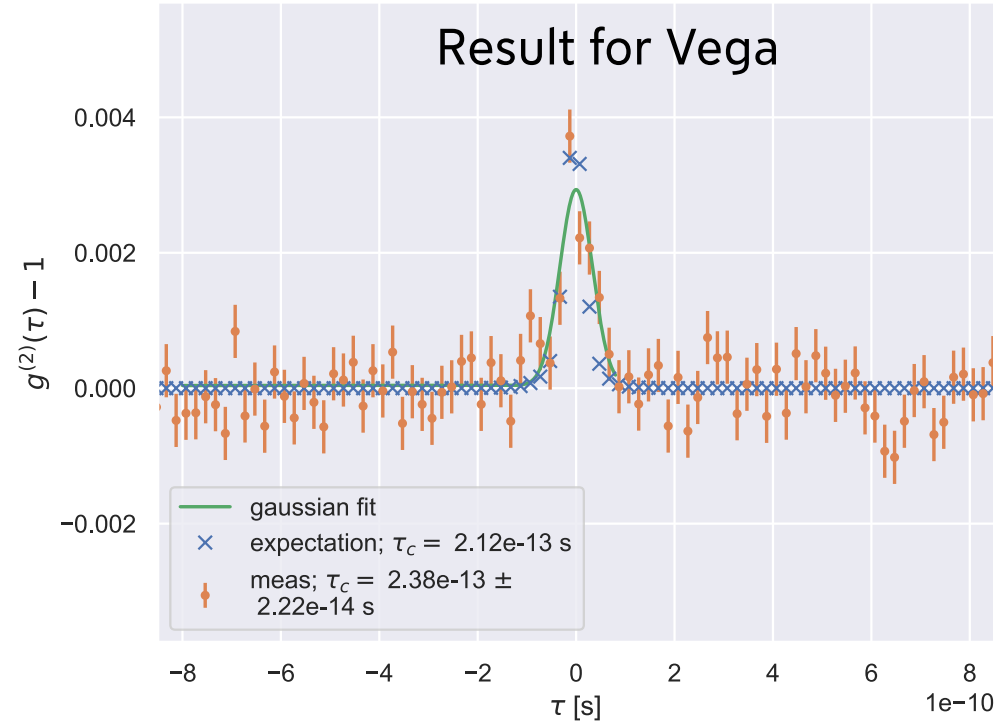


# Temporal Correlations at C2PU

## Observation Time



## App. Magnitude

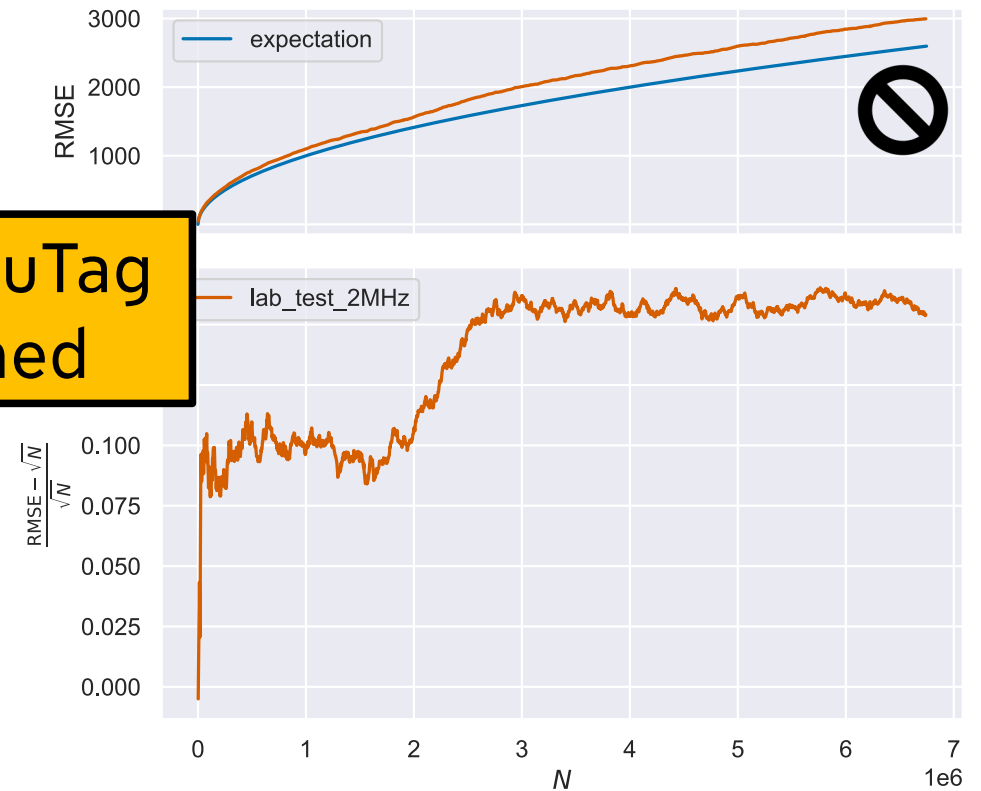
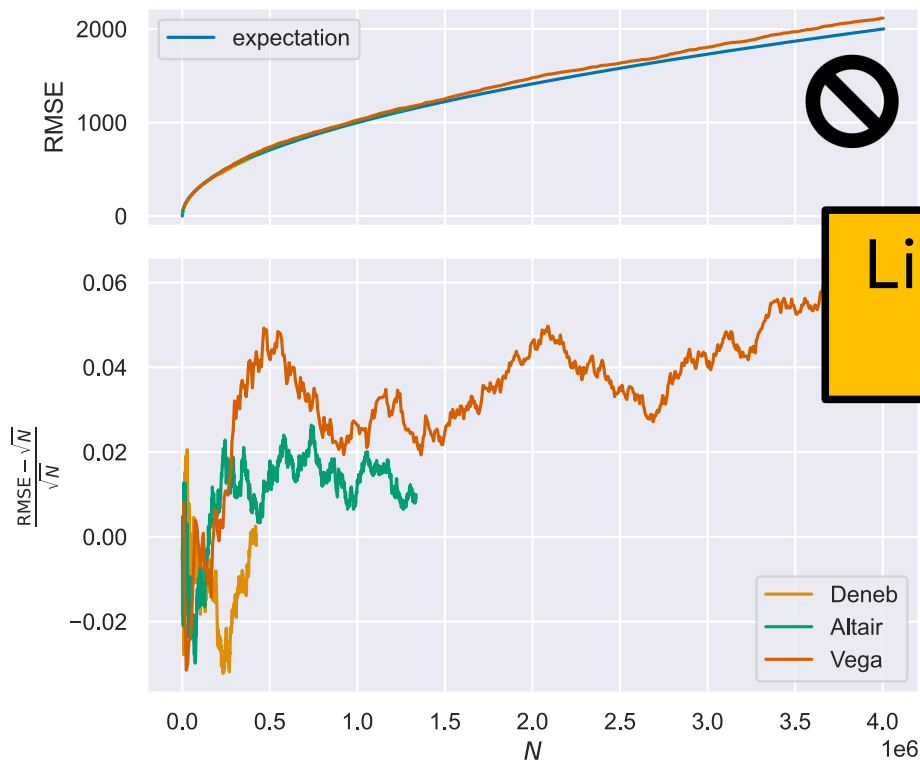


Star	Meas. $\tau_c$ [ps]	SNR
Vega	$0.238 \pm 0.0222$	10.72
Altair	$0.217 \pm 0.0350$	6.20
Deneb	$0.159 \pm 0.0605$	2.63

# Limitations of the configuration

@Telescope

In the Lab

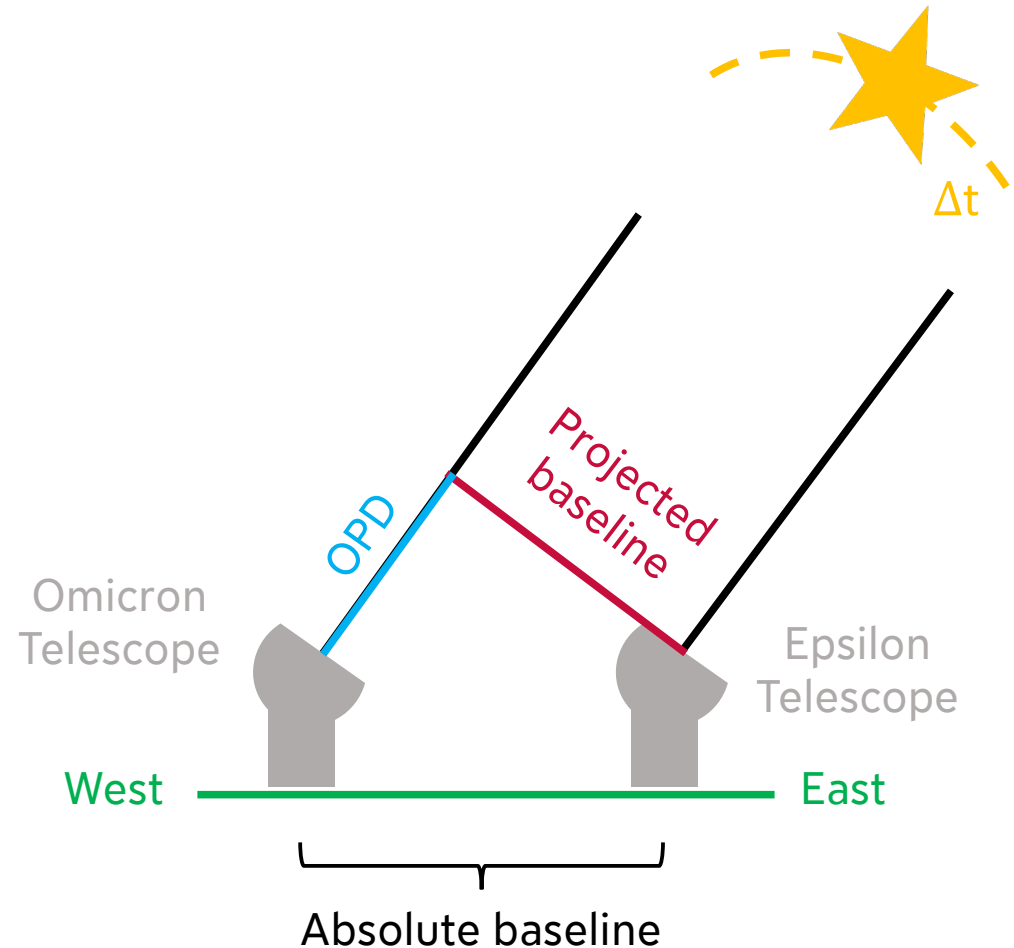


Limit of quTag is reached

# The Problem of moving Stars

# Calculation of baselines

- Absolute baseline of C2PU is 15m
- Baseline changes as star moves over the sky
- Calculate **projected baseline** for  $\Delta t$
- Calculate **optical path difference** (OPD) for  $\Delta t$
- Shift each histogram accordingly

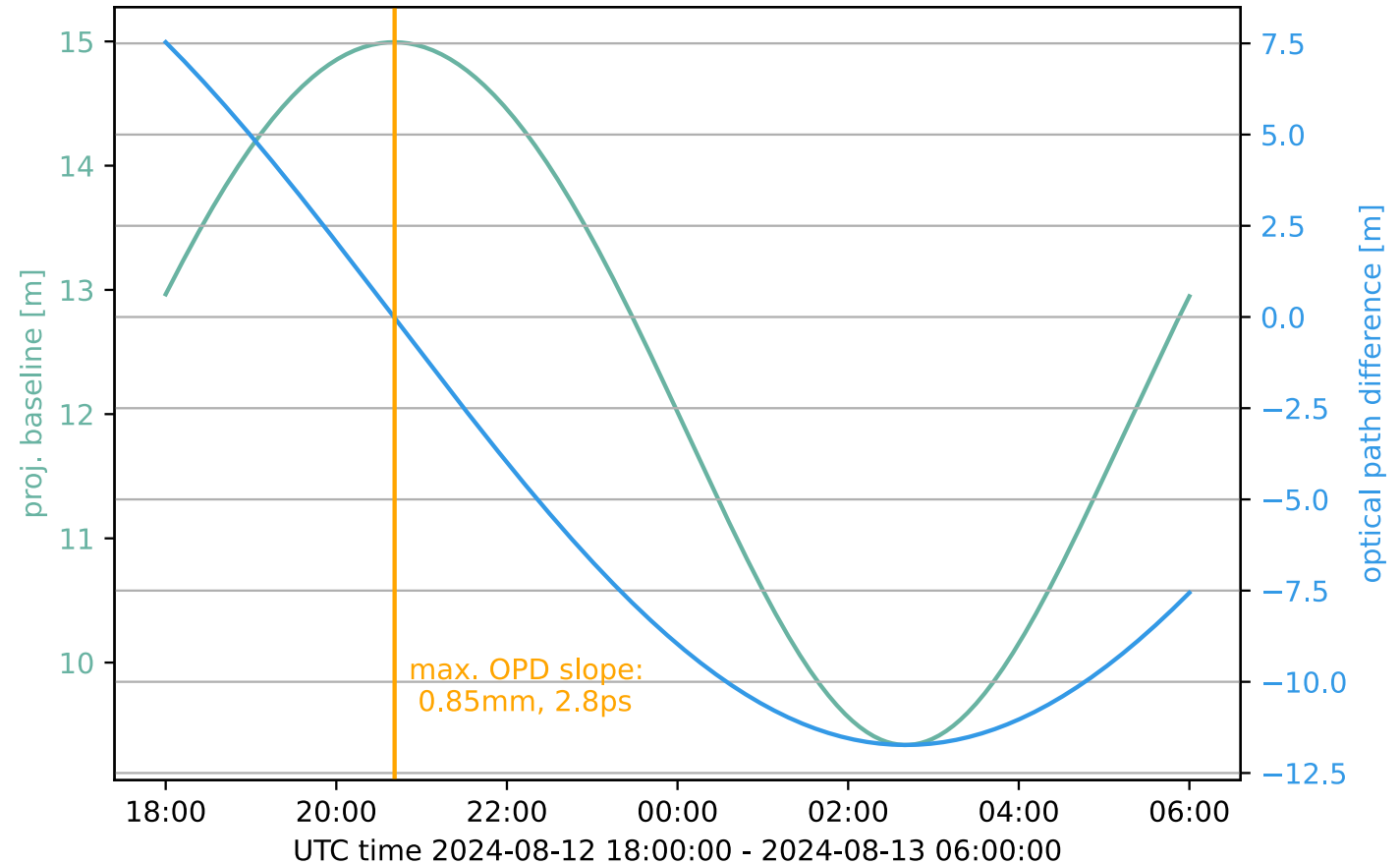




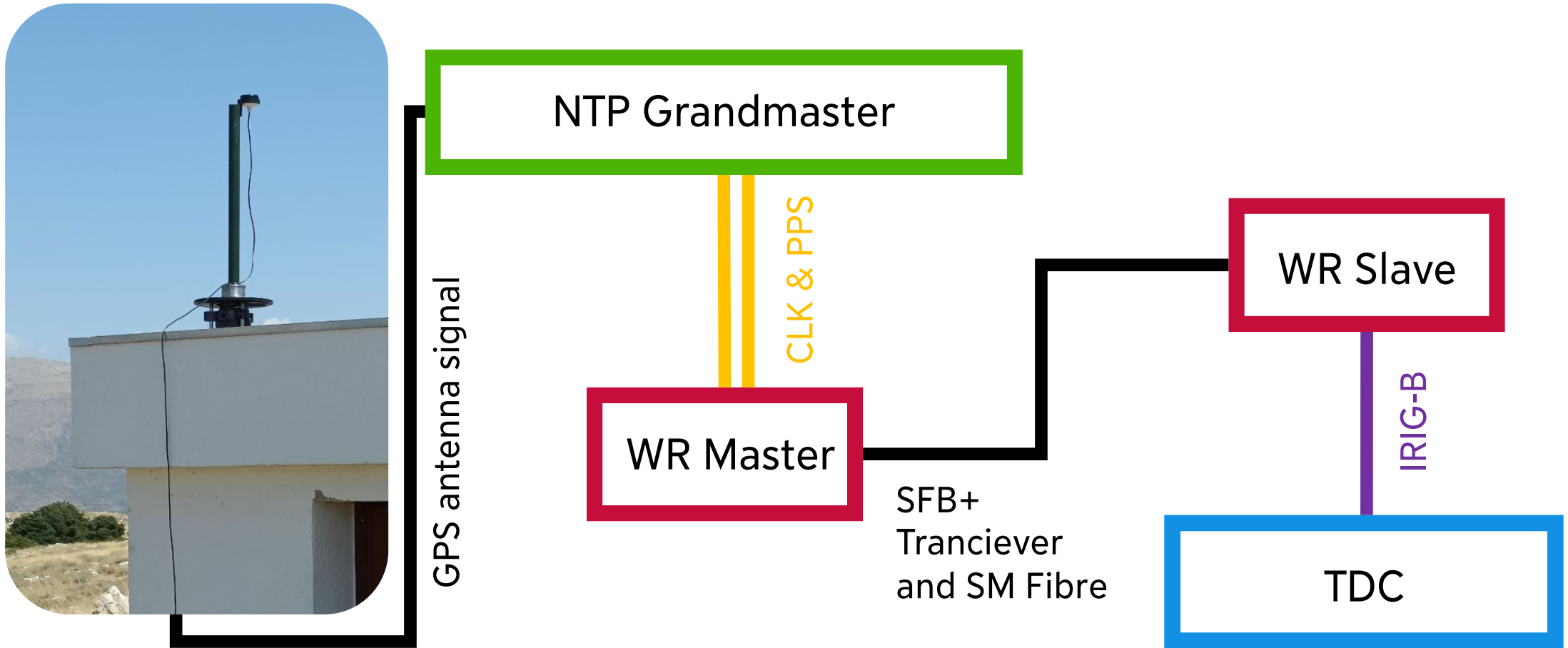
# Optical path difference (OPD)

Vega

- OPD slope needs to be  $<$  timing jitter of system
- Max. OPD slope: 0.85mm, 2.8ps per second
- Need UTC timestamps every second for OPD calculation



# Obtaining UTC timestamps



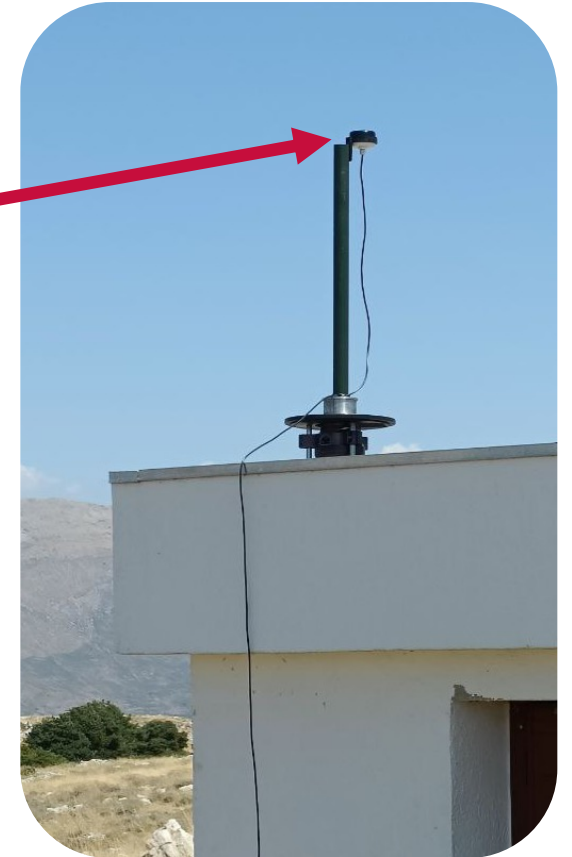
# Synchronization Hardware



NTP  
Grand-  
master

# NTP Grandmaster

- DTS 4160.Grandmaster from Bürk Mobatime
- 10MHz and PPS output to feed WR
- External time via GPS antenna
- GPS-synchronizable oscillator
- Oven Controlled Crystal Oscillator: 5 ps/s
- Provides UTC timestamps





# Synchronization Hardware

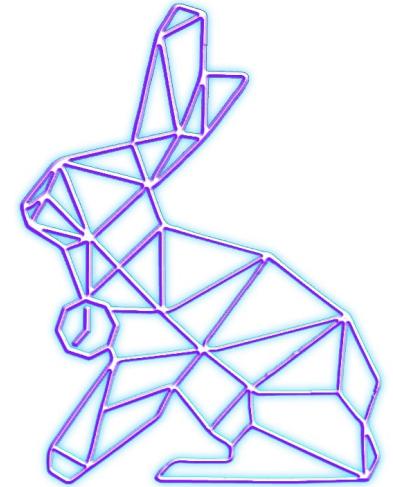
**White  
Rabbit  
LEN**



**NTP  
Grand-  
master**

# White Rabbit (WR)

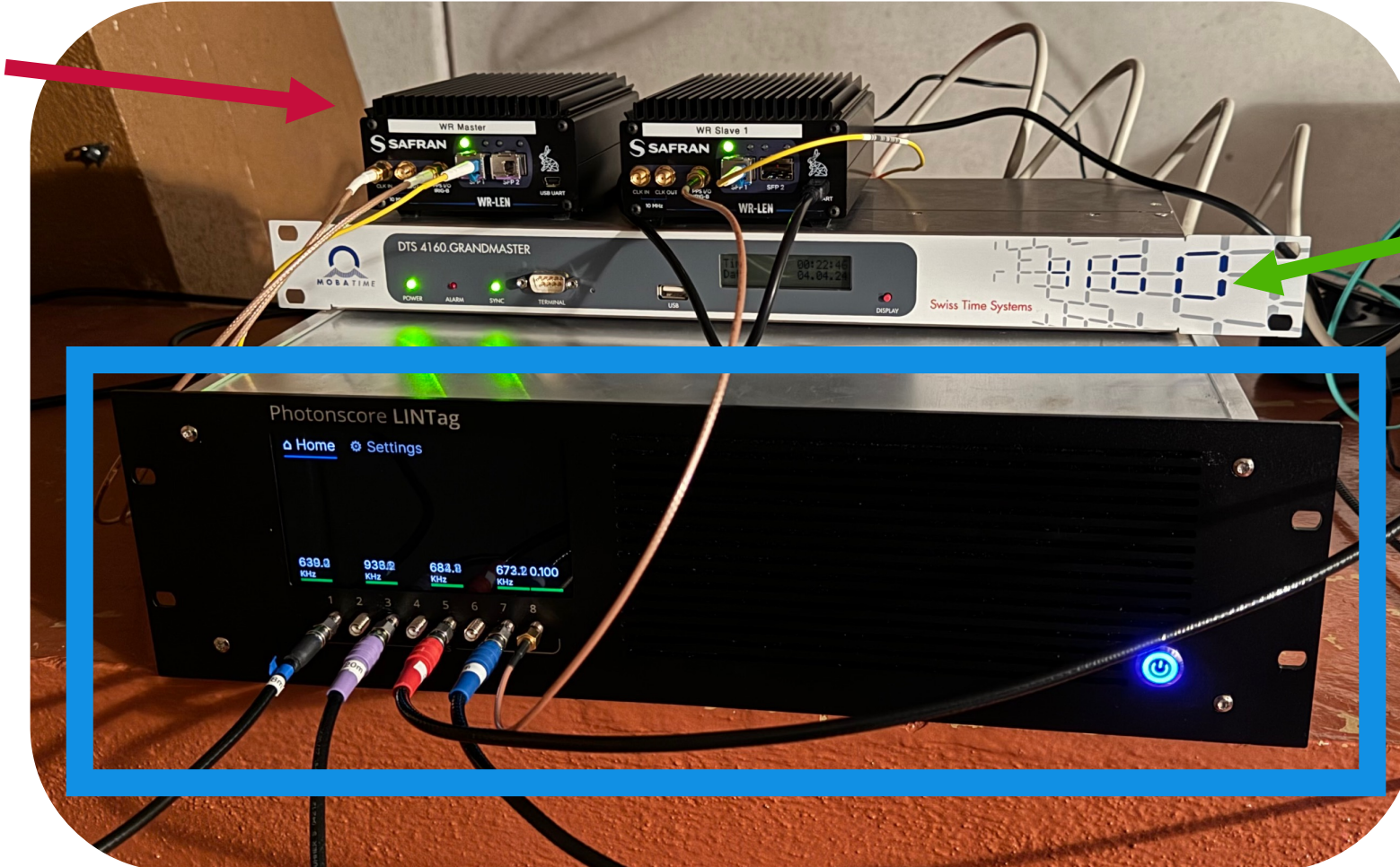
- WR LEN Standalone kit
- Synchronization over daisy chain configurations
- Distribute accurate timing
- Sub-nanosecond time accuracy
- Distance range: Over 80km fiber
- Dual optical fiber Ethernet 1G interfaces
- Synchronize IRIG-B device





# Synchronization Hardware

White  
Rabbit  
LEN



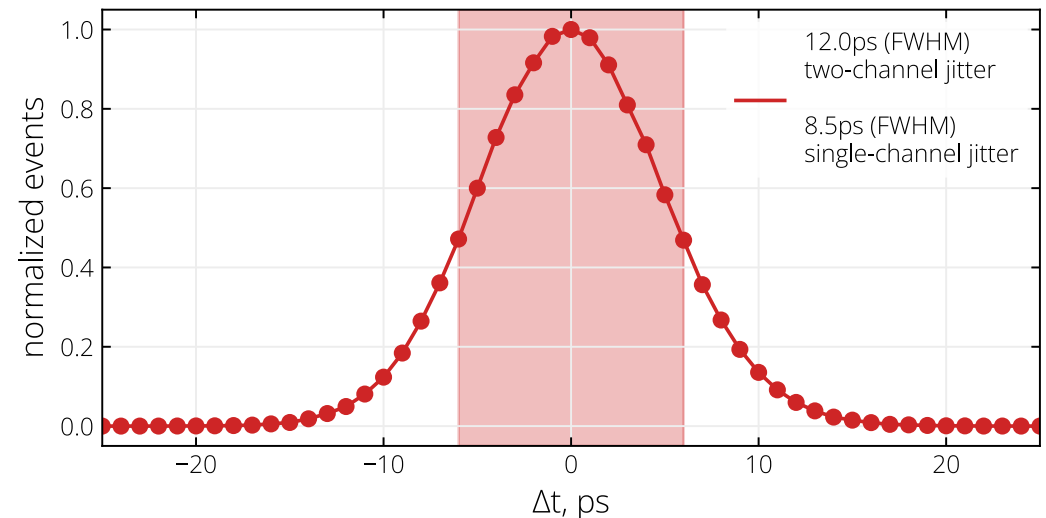
NTP  
Grand-  
master

LINTag  
TDC

# LINTag from Photonscore

- time-tagger system for ultra-fast data acquisition
- Temporal accuracy of 8.5ps (FWHM) / 3.6ps (RMS)
- 10G Ethernet SFB+ connection
- full-stack TCP/IP interface
- Transfer up to 400 MEvents/s
- Multi-Start Multi-Stop correlation

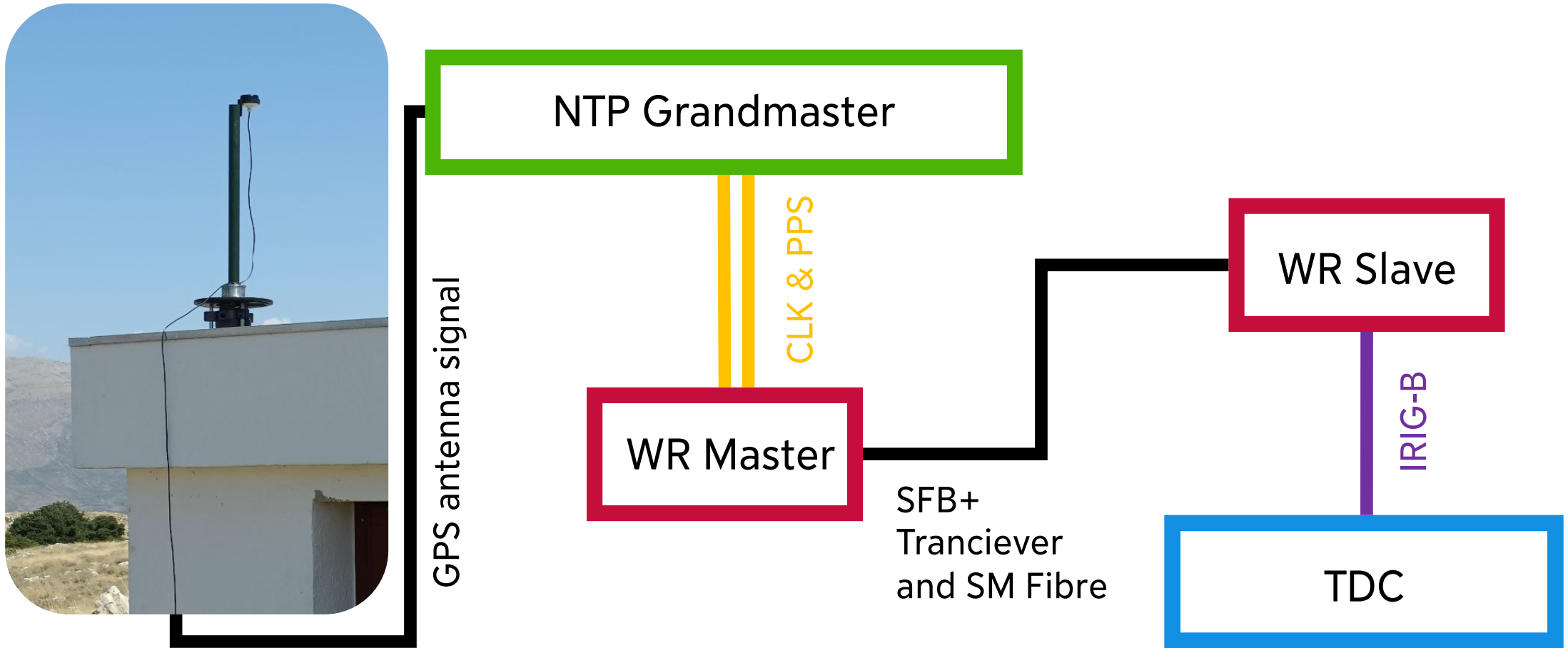
Timing jitter



**Photonscore.**  
PHOTON COUNTING MADE EASY

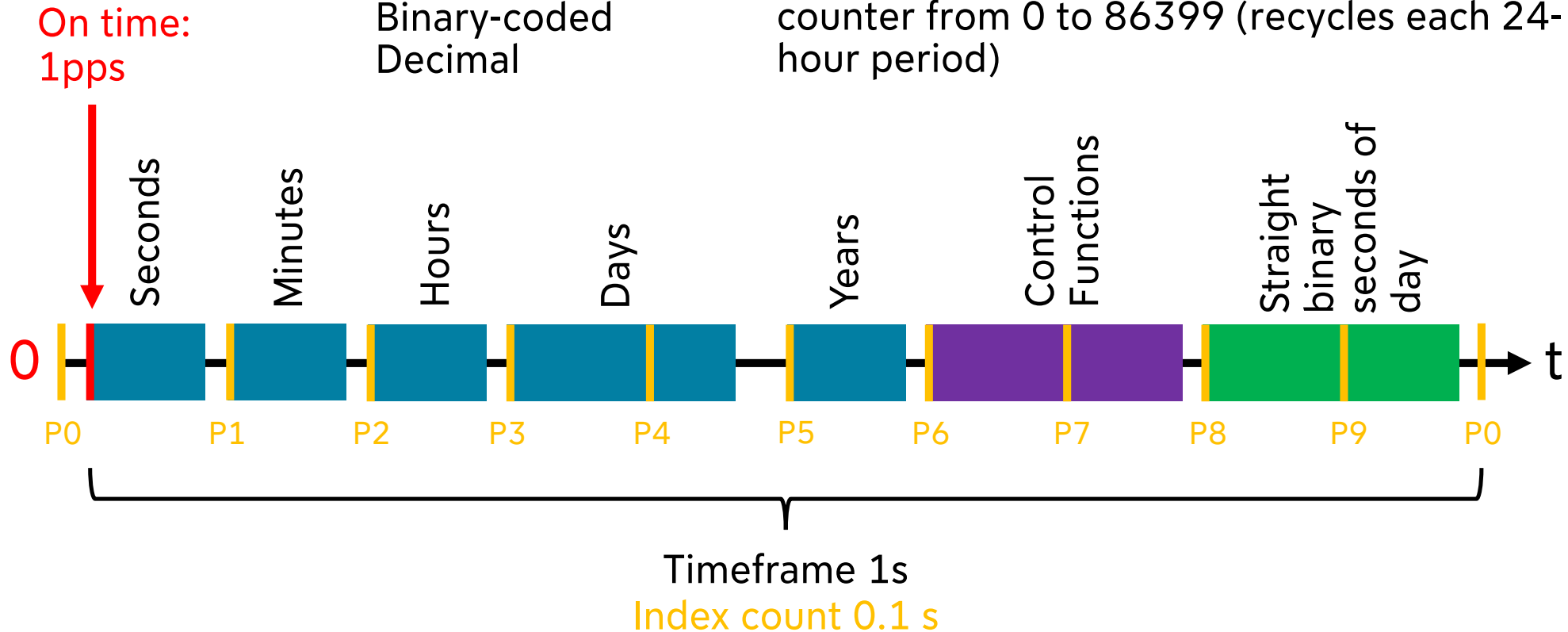


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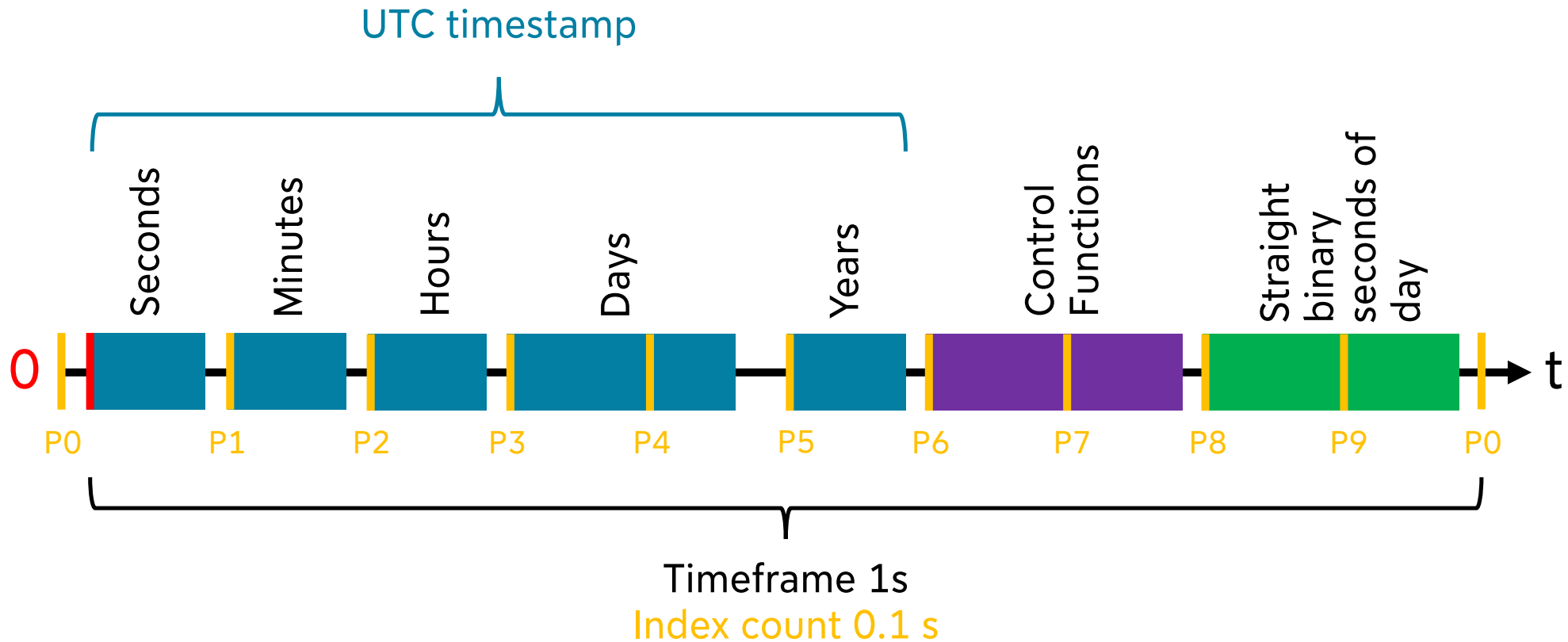


# IRIG-B timecode

- 100MHz bit rate
- Encoding type: Binary-coded Decimal
- Years as number 00-99 (00=2000)
- Straight binary seconds: 17 bit binary counter from 0 to 86399 (recycles each 24-hour period)



# IRIG-B timecode

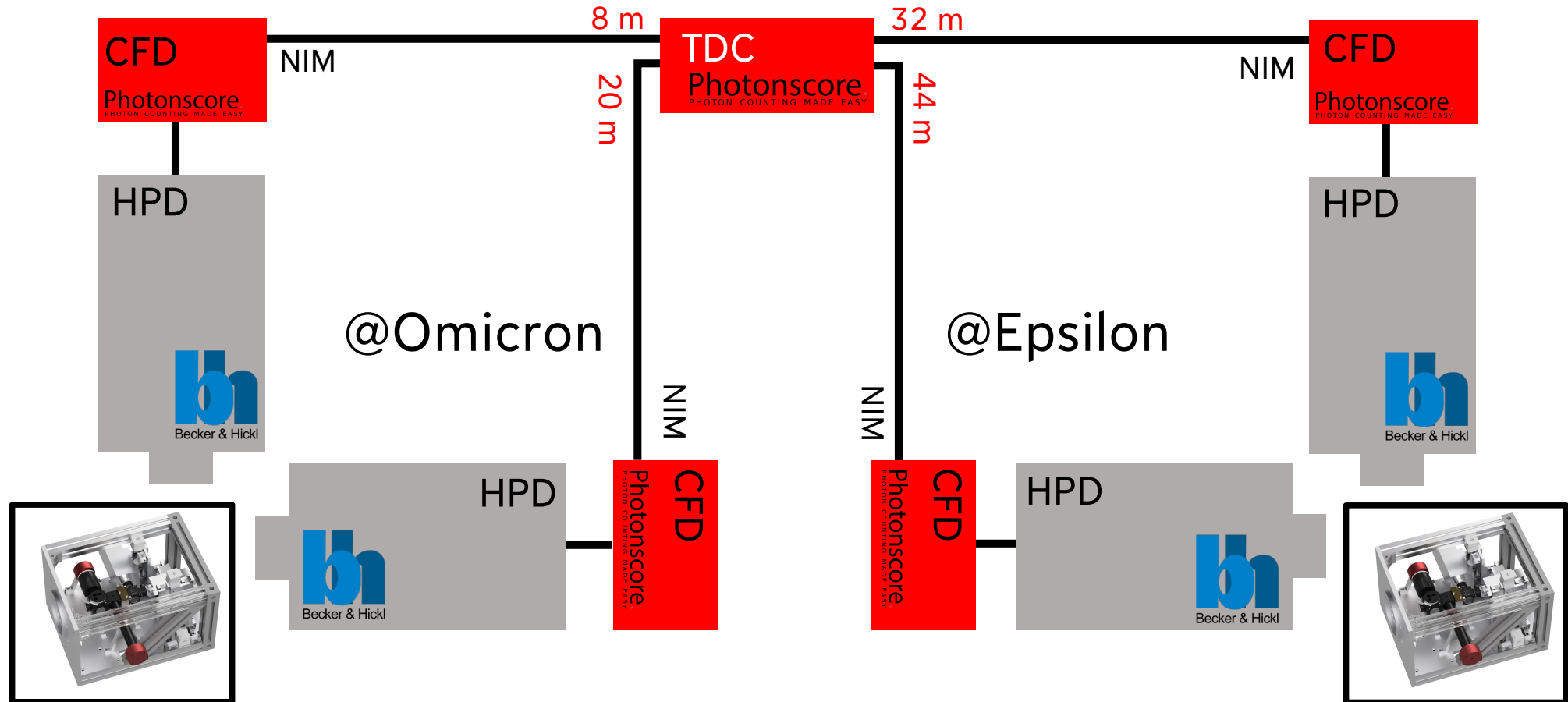


# Spatial Correlations using HPDs



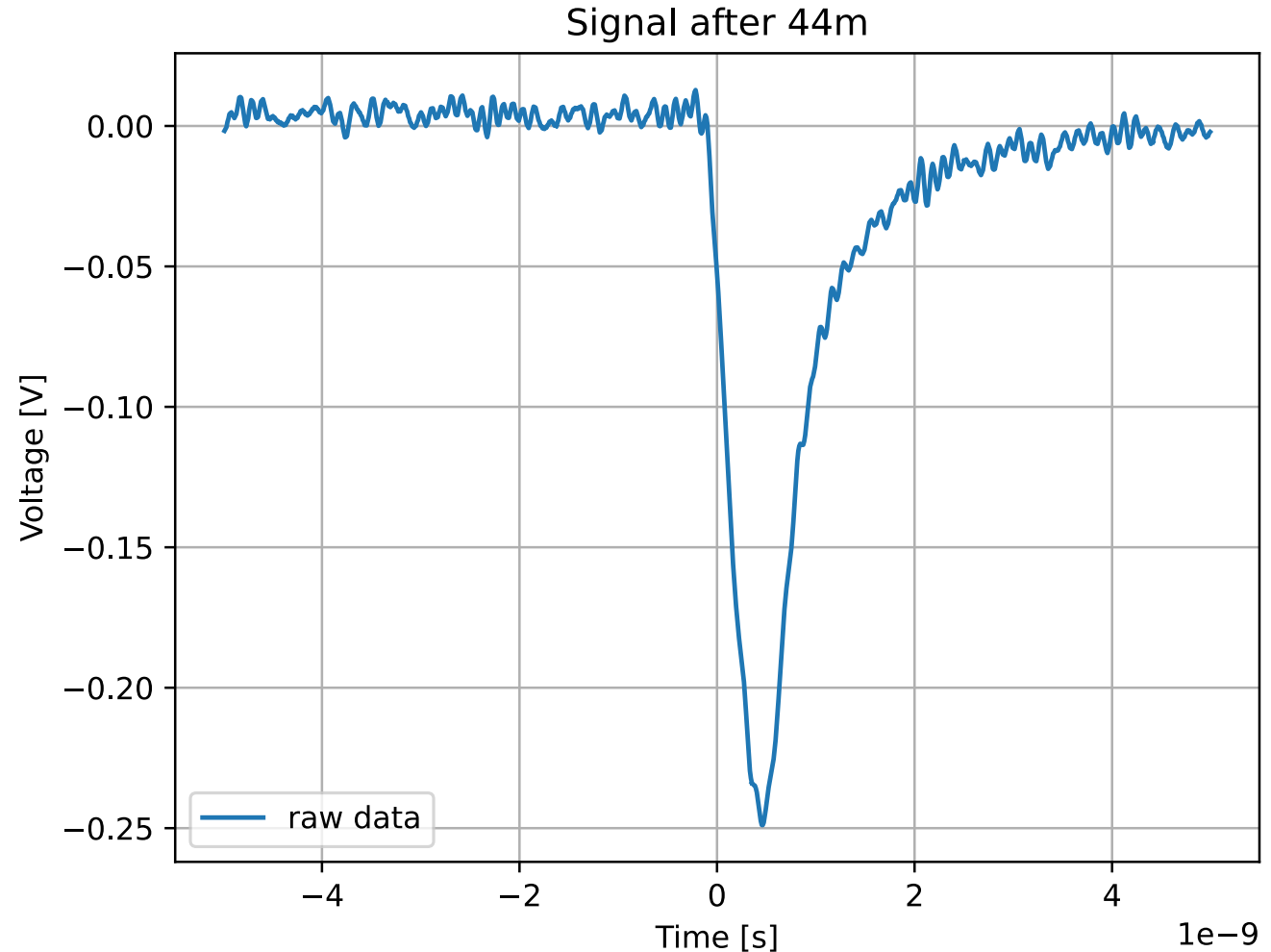


# Measurement Configuration



# Cabel lengths

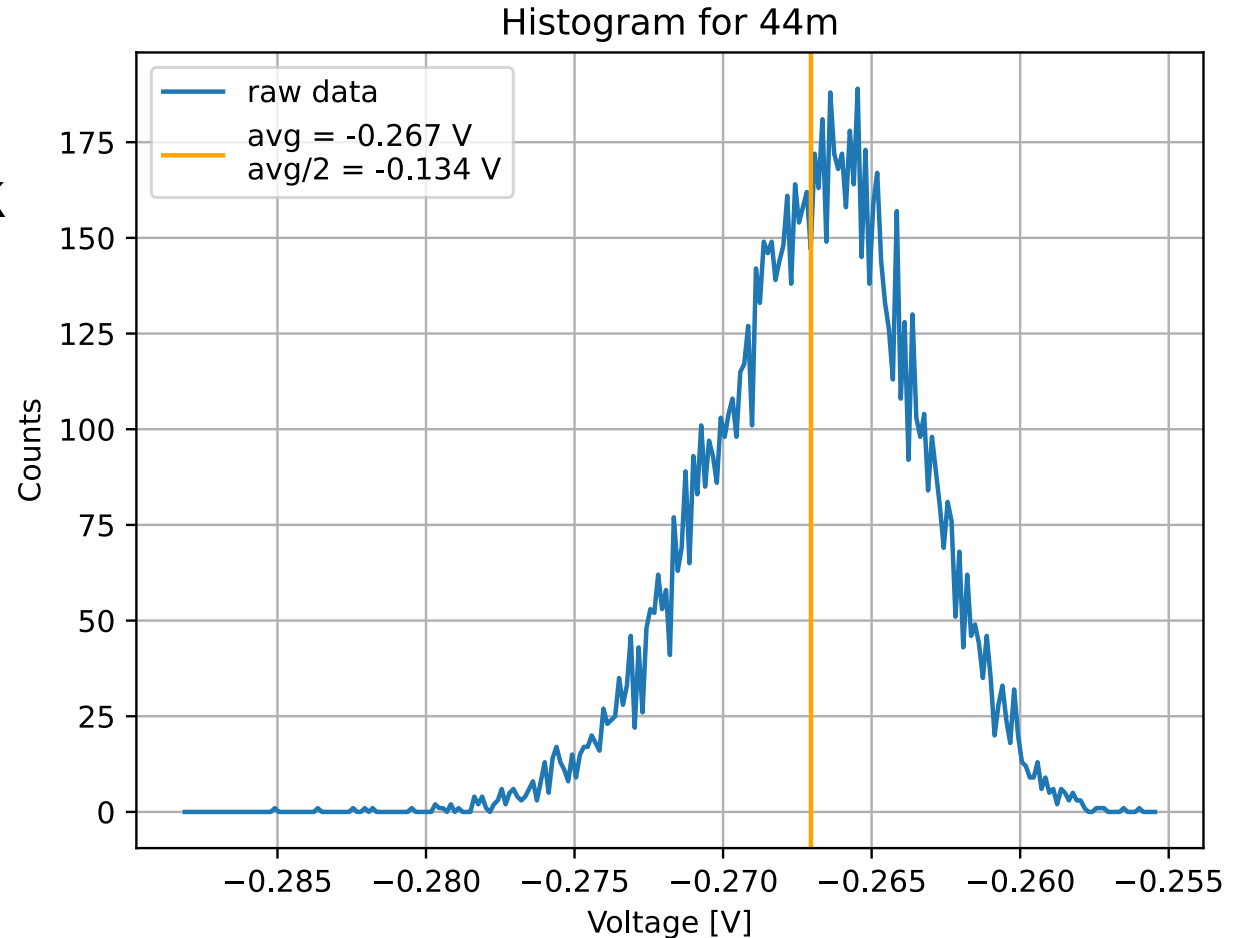
- Cable length difference: 12m to be far away from cross talk
- Hubert + Suhner 18GHz cable, semi-rigid
- NIM signal from CFD gets attenuated in cable



# Cabel lengths

- Cable length difference: 12m to be far away from cross talk
- NIM signal from CFD gets attenuated in cable
- Adjust trigger level in TDC

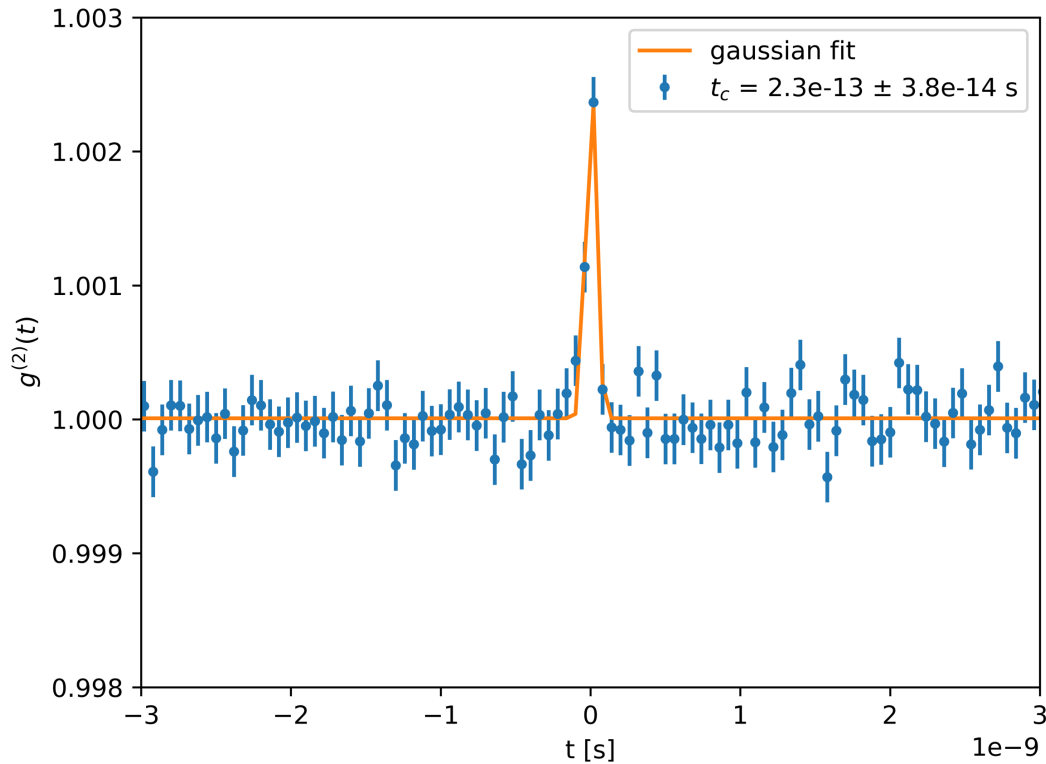
Cable length [m]	Trigger level [V]
8	-0.260
20	-0.211
32	-0.175
44	-0.134



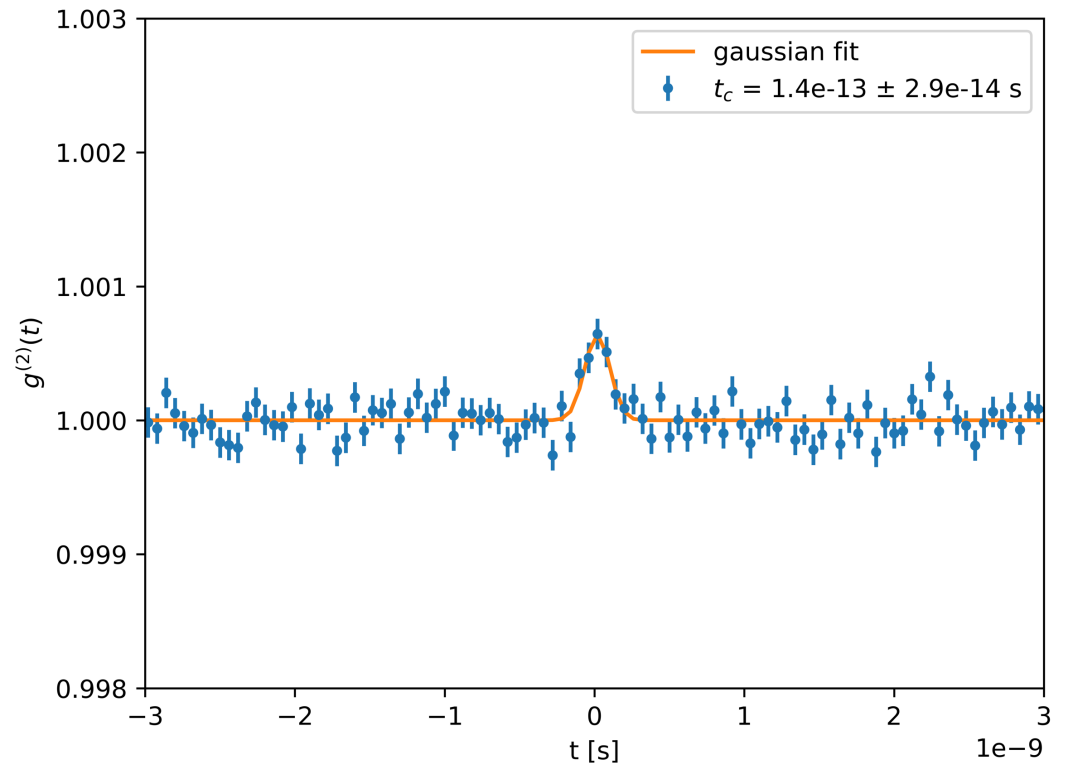
# Results for Vega using HPDs



## Bunching

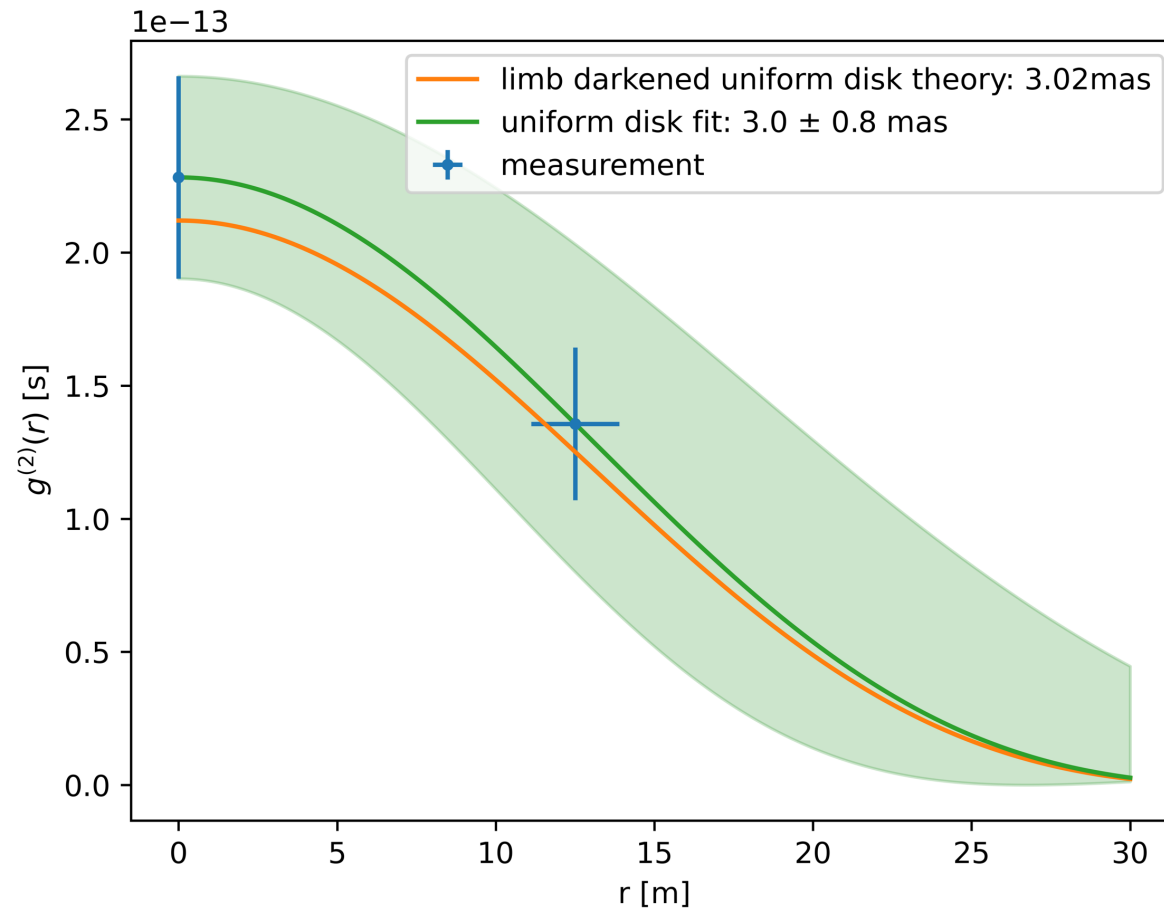


## Cross-Correlation





# Visibility curve for Vega

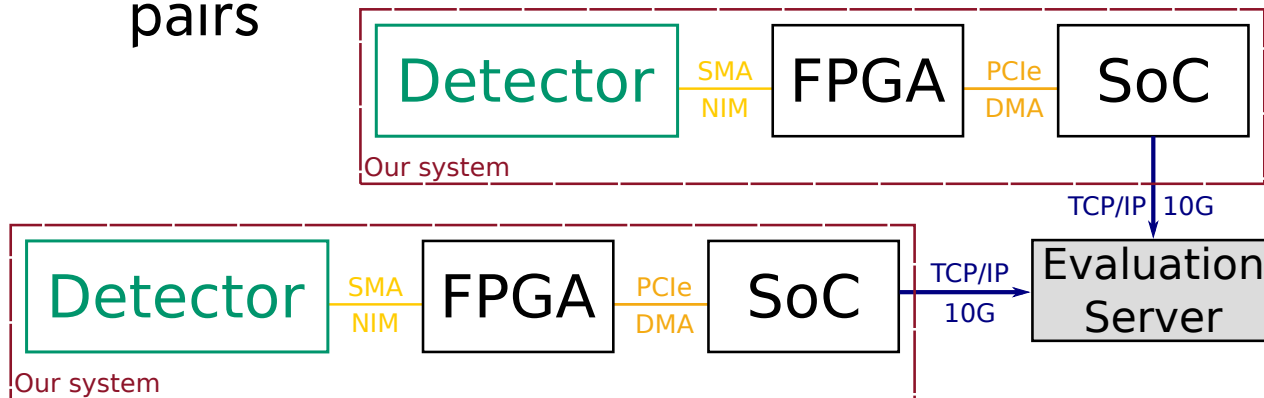
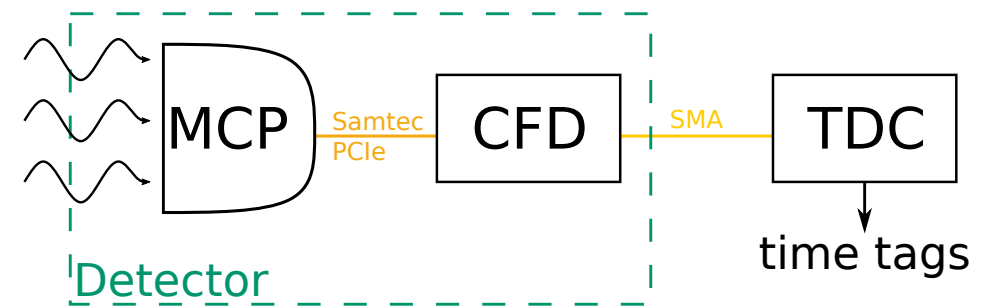


# High Throughput System (HTS)

# High Throughput System (HTS)

- Detector = LINPix and TDC = LINTag
- Timetags of FPGA will be sorted and compressed by SoC
- reliable, ordered and error-checked data delivery to server
- Evaluation server correlates all channel pairs

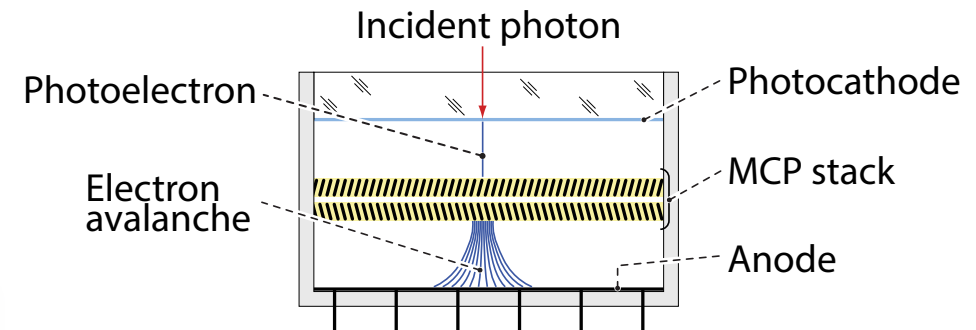
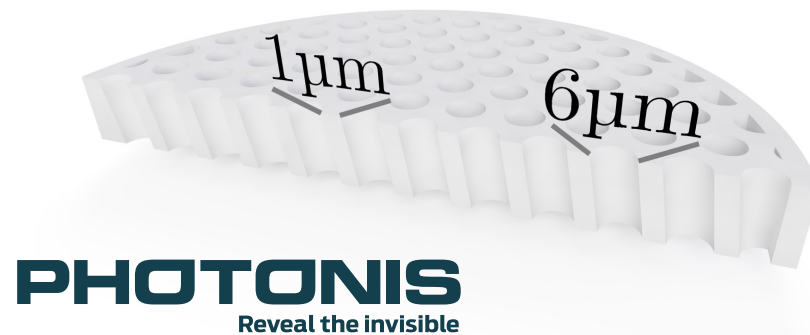
## LINPix from PhotonScore



- Integrated CFD
- NIM output  
→ long cables
- Count rates > 100MHz
- Ø8mm active area

# LINPix from Photonscore

## Ceramic Vacuum Assembly: Photocathode + Multichannel Plates



- Channel width =  $6\mu\text{m}$
- Channel distance =  $1\mu\text{m}$
- Channel tilt =  $5 - 10^\circ$
- Chevron stack  $\rightarrow$  gain  $> 10^6$



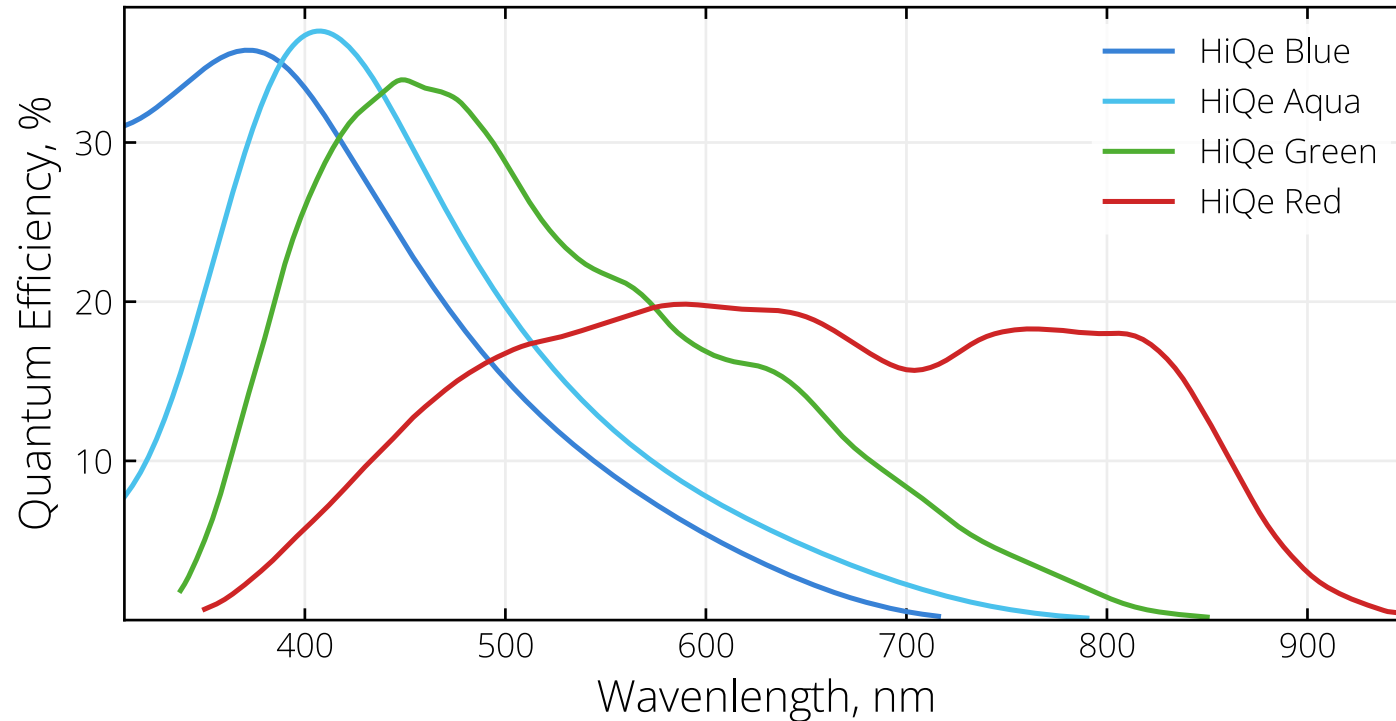
**Photonscore.**  
PHOTON COUNTING MADE EASY



# Possible Photocathodes for LINPix



Quantum efficiency



Used 3 HiQe Aqua and 1 HiQe Blue @405nm

# Spatial Correlations using HTS

2024

January							February							March						
S	M	T	W	T	F	S	S	M	T	W	T	F	S	S	M	T	W	T	F	S
	1	2	3	4	5	6				1	2	3				1	2			
7	8	9	10	11	12	13	4	5	6	7	8	9	10	3	4	5	6	7	8	9
14	15	16	17	18	19	20	11	12	13	14	15	16	17	10	11	12	13	14	15	16
21	22	23	24	25	26	27	18	19	20	21	22	23	24	17	18	19	20	21	22	23
28	29	30	31	25	26	27	28	29	24	25	26	27	28	29	30	31				

April							May							June						
S	M	T	W	T	F	S	S	M	T	W	T	F	S	S	M	T	W	T	F	S
	1	2	3	4	5	6				1	2	3	4							1
7	8	9	10	11	12	13	5	6	7	8	9	10	11	2	3	4	5	6	7	8
14	15	16	17	18	19	20	12	13	14	15	16	17	18	9	10	11	12	13	14	15
21	22	23	24	25	26	27	19	20	21	22	23	24	25	16	17	18	19	20	21	22
28	29	30	26	27	28	29	30	31	23	24	25	26	27	28	29					

July							August						
S	M	T	W	T	F	S	S	M	T	W	T	F	S
	1	2	3	4	5	6				1	2		
7	8	9	10	11	12	13	4	5	6	7	8	9	
14	15	16	17	18	19	20	11	12	13	14	15	16	17
21	22	23	24	25	26	27	18	19	20	21	22	23	24
28	29	30	31	25	26	27	28	29	30	31			

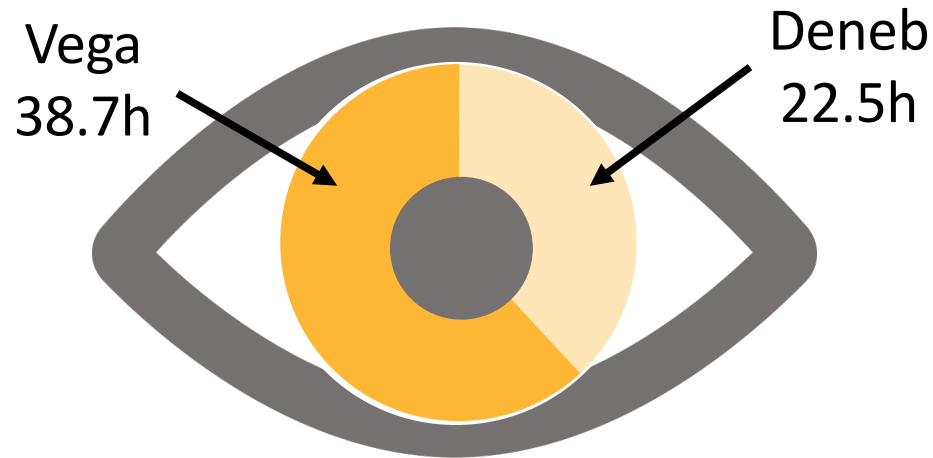
October							November							December							
S	M	T	W	T	F	S	S	M	T	W	T	F	S	S	M	T	W	T	F	S	
						1							1	2							1
6	7	8	9	10	11	12	3	4	5	6	7	8	9	8	9	10	11	12	13	14	
13	14	15	16	17	18	19	10	11	12	13	14	15	16	15	16	17	18	19	20	21	
20	21	22	23	24	25	26	17	18	19	20	21	22	23	22	23	24	25	26	27	28	
27	28	29	30	31	24	25	26	27	28	29	30	29	30	31							



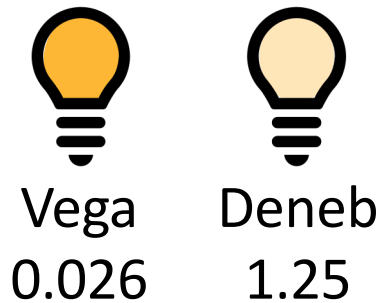


# Temporal Correlations using HTS

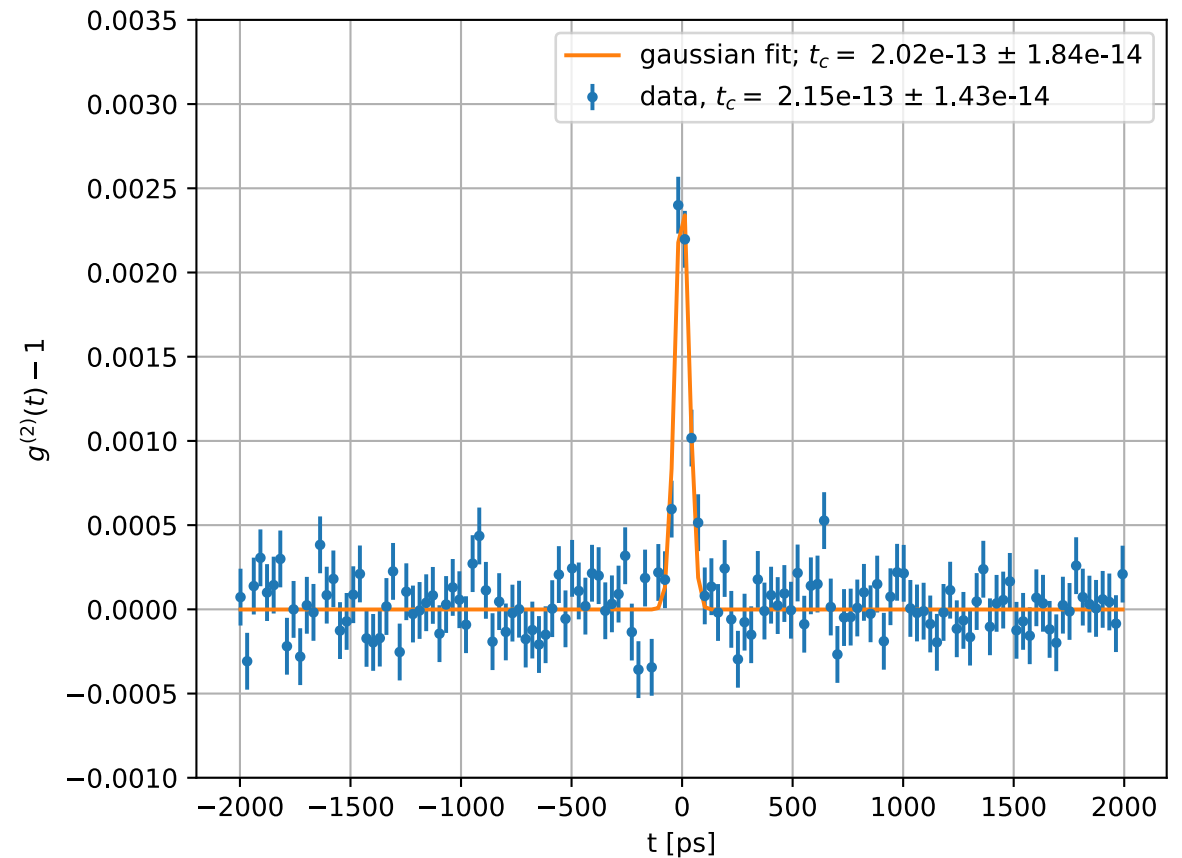
## Observation Time



## App. Magnitude



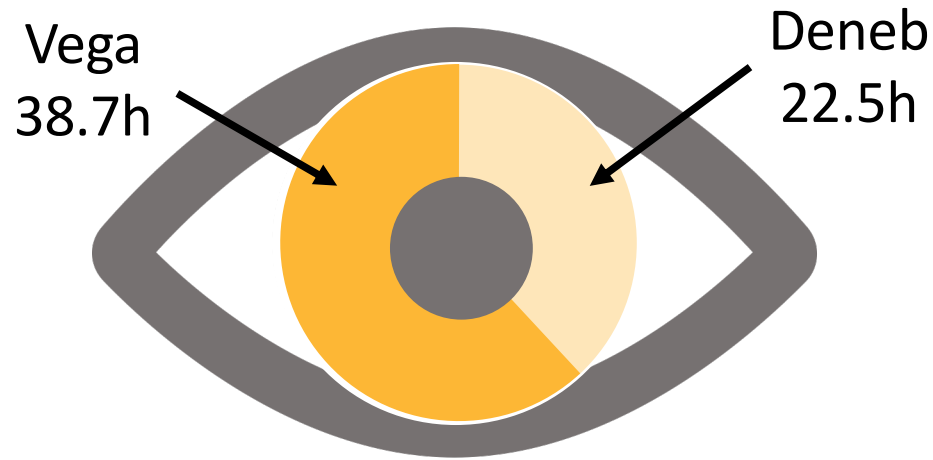
## Bunching for Vega



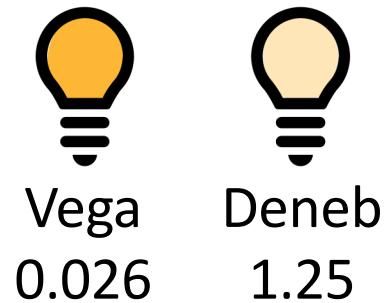


# Spatial Correlations using HTS

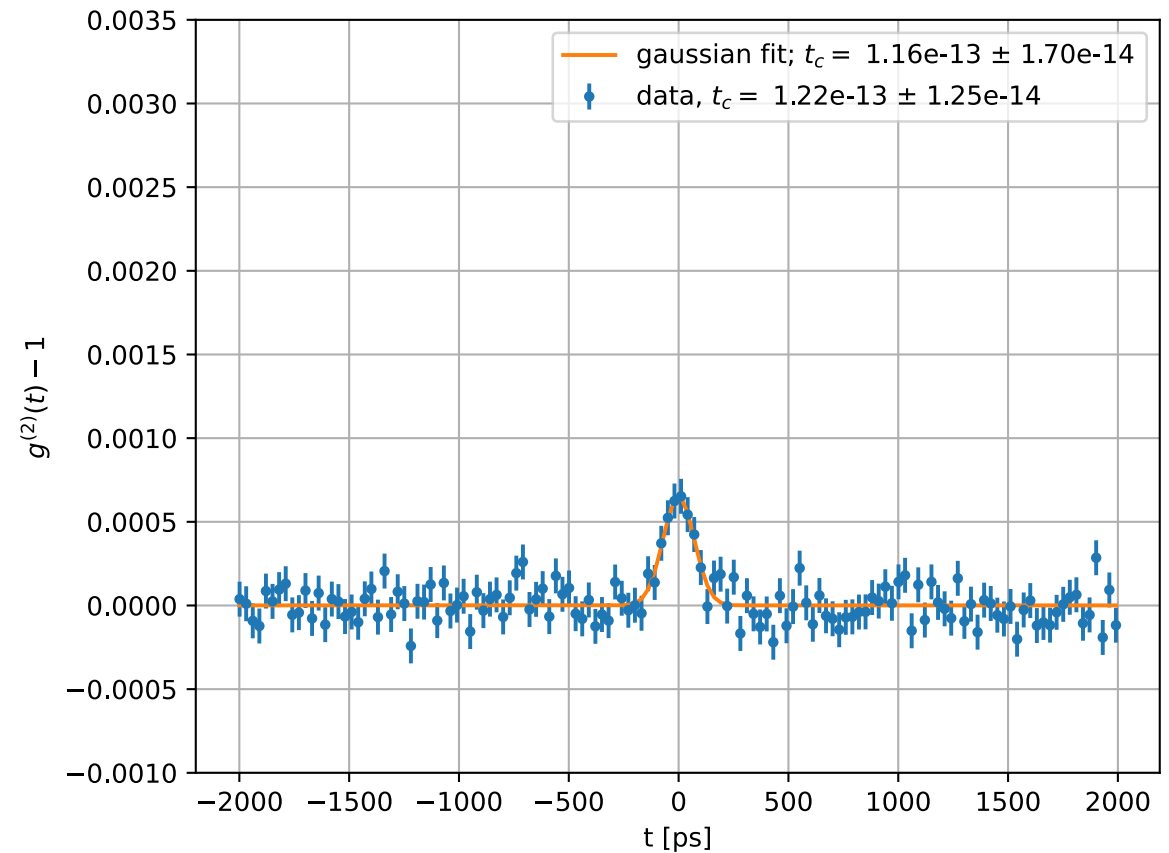
## Observation Time



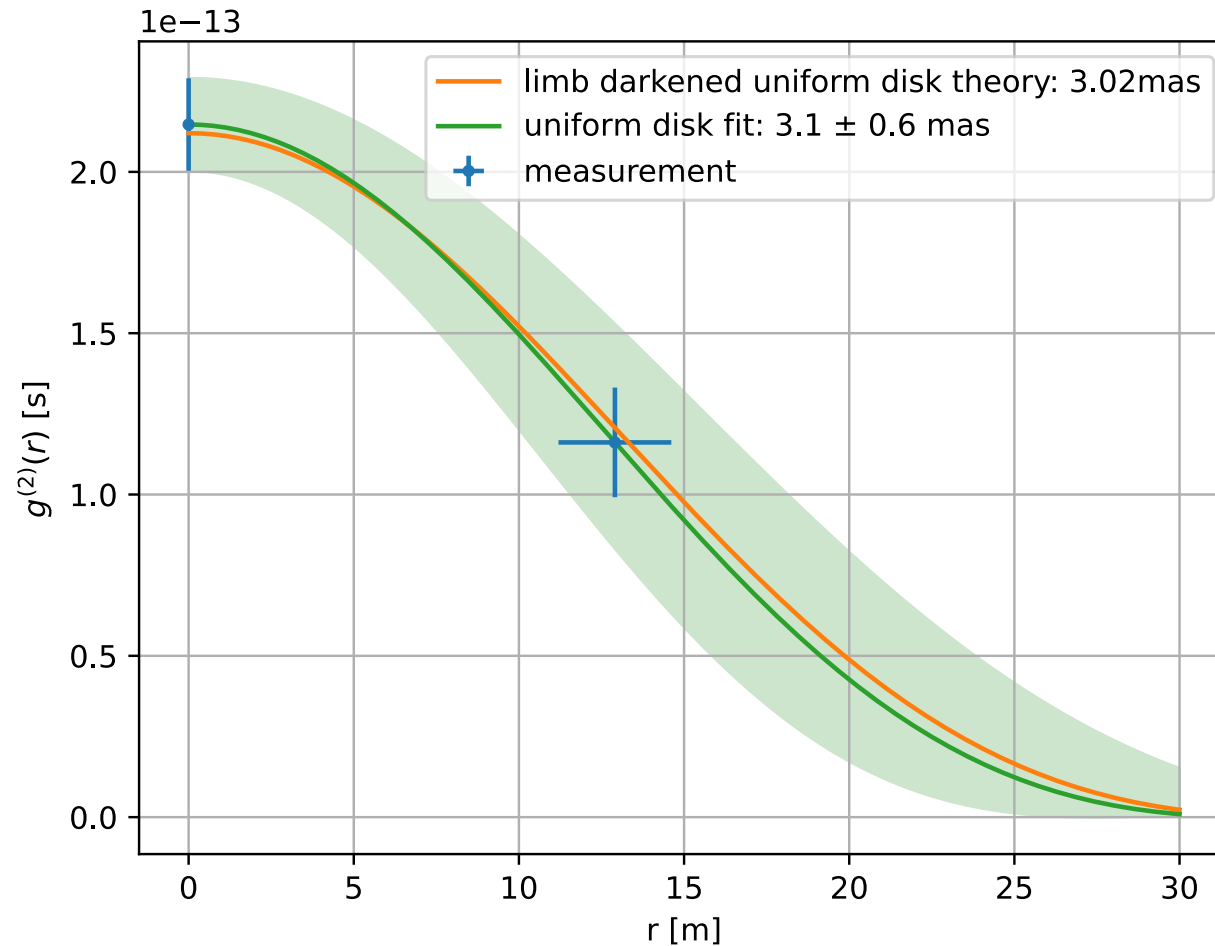
## App. Magnitude



## Result for Vega



# Visibility curve for Vega



# Summary & Outlook

# Summary and Outlook

## Summary

- Verified stability of setup
- Measured Bunching and spatial correlations using HPDs
- Successfully tested LINPix and LINTag



To Do:

- 1) Check filter position
- 2) Higher count rates?
- 3) Bursts?
- 4) Km baselines?



## Outlook

- 1) Measure backreflection
- 2) Try 100MHz bunching
- 3) Check burst capability of LINPix
- 4) Try WR with long fibers and test synchronizability of multiple TDCs





**Thank you  
for your attention**

# Image References



- HPD: taken from Becker&Hickl manual
- CFD1: [https://upload.wikimedia.org/wikipedia/commons/8/8f/Constant\\_fraction\\_1.svg](https://upload.wikimedia.org/wikipedia/commons/8/8f/Constant_fraction_1.svg)
- CFD2: [http://lmu.web.psi.ch/docu/manuals/bulk\\_manuals/software/TDC/CFD\\_signals.png](http://lmu.web.psi.ch/docu/manuals/bulk_manuals/software/TDC/CFD_signals.png)
- quTAG: taken from quTools website
- WR: <https://www.eenewseurope.com/en/white-rabbit-deal-boots-timing-synchronisation/>
- WR Len: taken from Safran website
- NTP Grandmaster: taken from Bürk Mobatime website
- LINTag plot and image: taken from Photonscore manual and website
- Vacuum assembly: copyright Photonis
- MCP: copyright Stefan Richter
- MCP schematic: copyright Yury Prokazov
- LINPix: taken from Photonscore website
- Quantum efficiency: taken from LINPix Photonscore datasheet